

# Chirally symmetric droplets in the core of magnetars

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## 1 Motivation: compact stars

- Compact stars: what's the EoS of strongly interacting matter for low temperatures and high densities?
- The order, strength and location of the strong interactions phase transitions: crucial to establish new classes of compact stars [1]
- Magnetars [2]:  $B \sim 10^{19} - 10^{20} \text{G}$ !  $\Rightarrow$  How is the phase structure altered?

**Crucial: what are the time scales involved?**

## 2 Building the effective theory

### 2.1 The need of an effective theory description

- Surface tension,  $\Sigma$ : energy per unit area needed to create a bubble  $\Rightarrow$  Key quantity!
- Ideally: compute  $\Sigma$  in the high  $\mu$  low  $T$  region of the QCD phase diagram
- Reality: first principle approaches are difficult to apply in this region
  - Lattice QCD: sign problem
  - pQCD: phenomenologically interesting values of  $\mu$  are already too low, 400 – 500 MeV

**Solution: Estimating  $\Sigma$  in an effective theory**

### 2.2 General Framework

- We describe the transition in a Chiral effective theory, The Linear Sigma Model coupled to quarks (LSM<sub>q</sub>) [3]. The Lagrangian is given by:
 
$$\mathcal{L} = \bar{\psi}_f [i\gamma^\mu \partial_\mu - g(\sigma + i\gamma_5 \boldsymbol{\tau} \cdot \boldsymbol{\pi})] \psi_f + \frac{1}{2}(\partial_\mu \sigma \partial^\mu \sigma + \partial_\mu \boldsymbol{\pi} \cdot \partial^\mu \boldsymbol{\pi}) - \frac{\lambda}{4}(\sigma^2 + \boldsymbol{\pi}^2 - v^2)^2 + h\sigma$$
- Both QCD with two massless quark flavors and the  $O(4)$  LSM<sub>q</sub> belong to the same universality class [4]
- The model reproduces the low energy phenomenology of strong interactions  $\Rightarrow$  Parameter fixing
- Spontaneous and (small) explicit break of Chiral symmetry are contained in the meson self interaction
- Coupling the system with an external magnetic field only affects the quark sector,  $\partial_\mu \rightarrow \partial_\mu + ieA_\mu$
- Phase conversion is well described in the absence of pions [5].

### 2.3 Effective potential at one loop

- Quark gas: thermal bath for the long-wavelength  $\sigma$  field
- Effective potential: integrating over quarks, treating  $\sigma$  classically
- We take the cold and dense approximation, i.e.  $T = 0$  and finite quark chemical potential:

$$V_{\text{eff}}(\bar{\sigma}) = U_{\text{cl}}(\bar{\sigma}) + U_f^{\text{vac}}(\bar{\sigma}, B) + U_f^{\text{med}}(\bar{\sigma}, \mu, B)$$

where  $U_f^{\text{vac}}$  denotes the fermionic vacuum contribution,

$$U_f^{\text{vac}} = -\frac{N_c}{2\pi^2} \sum_f (q_f B)^2 \left[ \zeta_H'(-1, x_f) + \frac{x_f^2 - x_f}{2} \log x_f + \frac{x_f^2}{4} \right],$$

and  $U_f^{\text{med}}(\bar{\sigma}, \mu, B)$ , the medium contribution,

$$U_f^{\text{med}} = -\frac{N_c}{4\pi^2} \sum_f \sum_{\nu=0}^{\nu_{\text{max}}} (2 - \delta_{\nu 0}) |q_f| B \left[ \mu \sqrt{\mu^2 - M_{fB}^2} + M_{fB}^2 \log \left( \frac{\mu + \sqrt{\mu^2 - M_{fB}^2}}{M_{fB}} \right) \right].$$

## 3 Surface tension and nucleation

- We consider homogeneous nucleation. Two possibilities:
  - Thermal activation
  - Quantum nucleation
- Physical setting [6]:  $T = 10 - 20 \text{ MeV}$
- Temperatures are high enough for thermal activation to take place. Quantum nucleation is negligible [1, 7, 8]
- For the chemical potential range we study, the cold approximation for the effective potential is still valid

### 3.1 Extracting nucleation parameters from the effective potential

- Our aim is to obtain the qualitative behavior, not numerical precision.
- We write the effective potential in the Ginzburg-Landau form, quartic fit [5]:

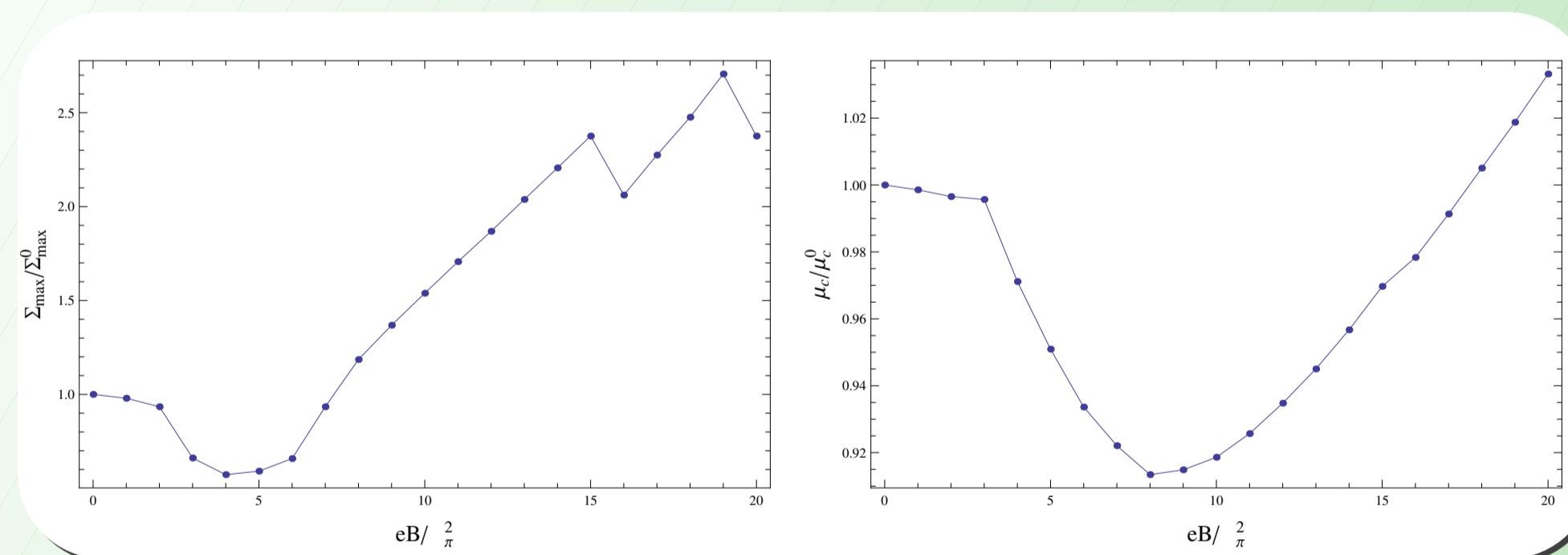
$$V_{\text{eff}} \approx \sum_{n=0}^{n=4} a_n \phi^n$$

- Although unable to reproduce all three minima, provides a good description in the region of interest for nucleation
- Simplifying limit: thin wall  $\Rightarrow$  nucleation parameters are given by analytic expressions of  $a_n$
- The procedure is valid for  $\mu$  in the range between  $\mu_c$  and  $\mu_{sp}$

## 4 Results

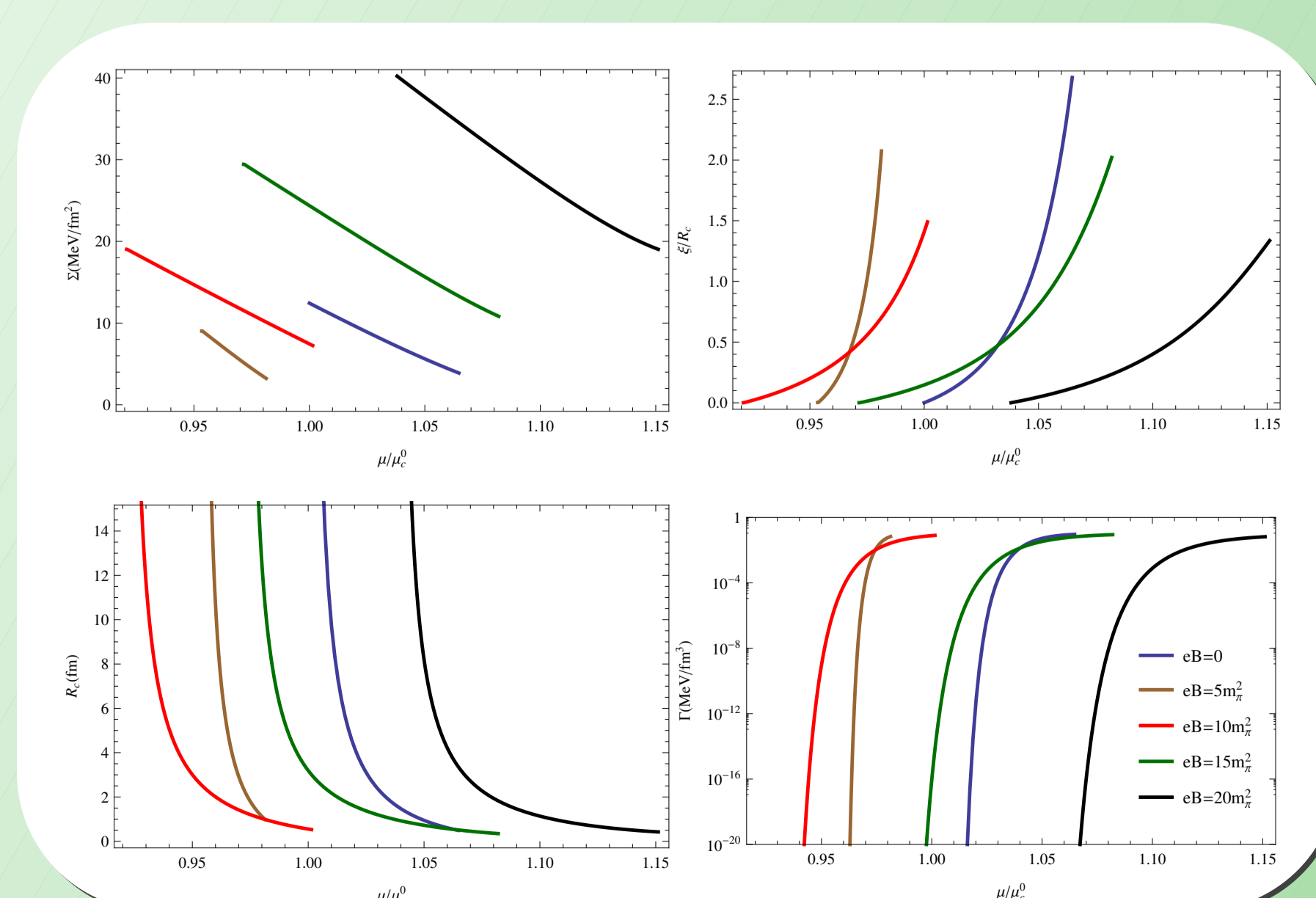
### 4.1 Landau level filling and oscillations

- Last occupied Landau level:  $\nu_{\text{max}} = \left\lfloor \frac{\mu^2 - g^2 \sigma^2}{2|q_f|B} \right\rfloor$
- Varying  $B$ , the number of filled Landau levels change  $\Rightarrow$  oscillations
- Similar to the ones found in the NJL model [9]



### 4.2 Nucleation parameters

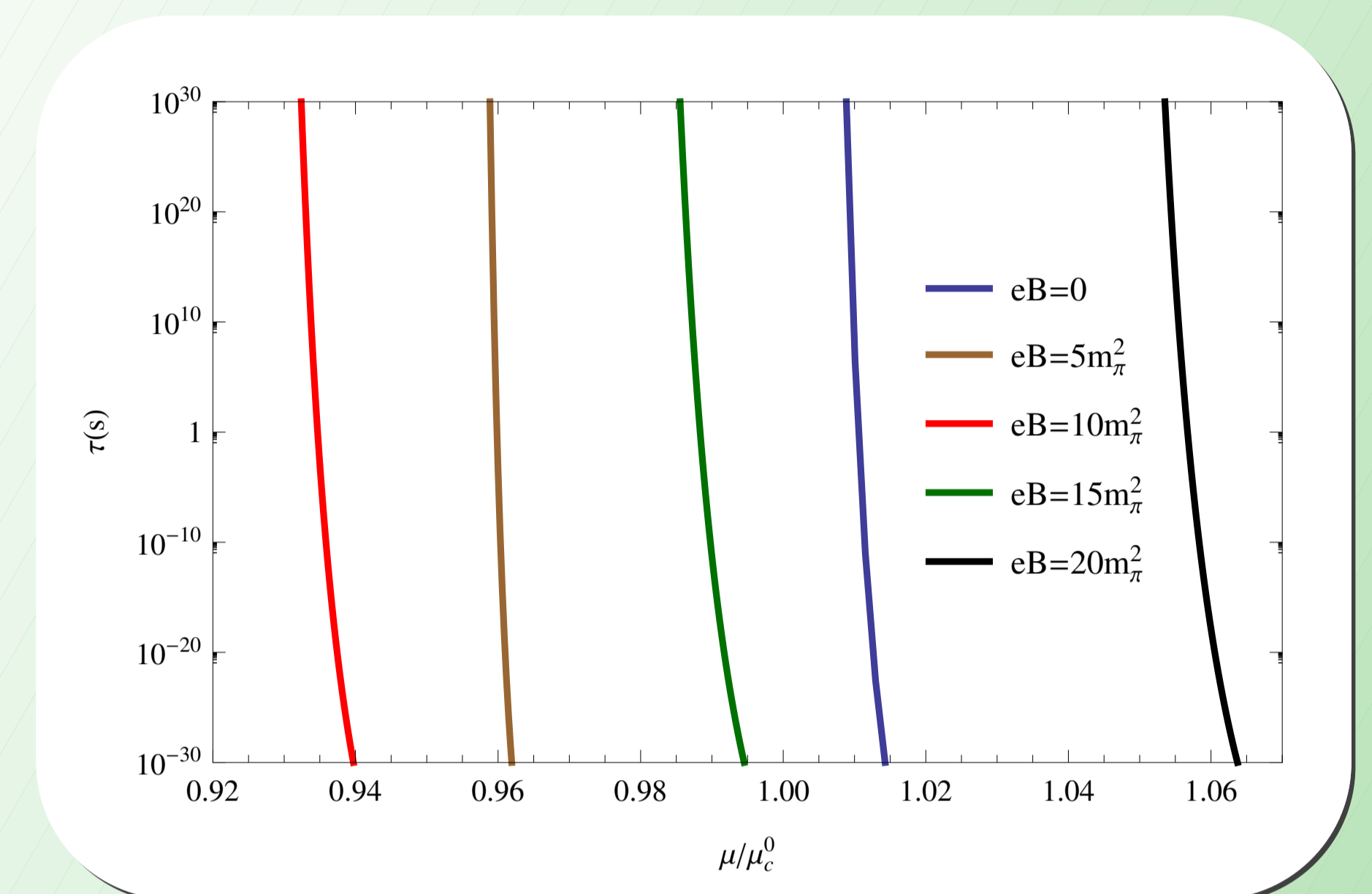
- Nucleation of chirally symmetric droplets
- Thin-wall is a good approximation for  $\xi/R_c$  small
- Nucleation rate:  $\Gamma \sim T_f^4 e^{-E_b/T_f}$ , for  $T_f = 30 \text{ MeV}$
- All plot lines begin at  $\mu_c$  and end at  $\mu_{sp}$



## 4.3 Nucleation time

- Extremely Important information!
- Time needed to nucleate a single critical bubble in a volume of  $1 \text{ km}^3$ :

$$\tau \equiv \left( \frac{1}{\text{km}^3} \right) \frac{1}{\Gamma}$$



## 5 Conclusion

- For magnetic fields up to  $5 m_\pi^2$ : decrease of the metastable region  $\Rightarrow$  phase conversion via spinodal explosion facilitated
- Non trivial competition between  $\Sigma$  and  $R_c$ : surface tension may increase and the nucleation time still decrease  $\Rightarrow$  **The surface tension alone is not able to give all the needed information about nucleation in magnetar matter!**
- Nucleation is favored for  $eB \approx 10 m_\pi^2$
- All these interesting features take place in the phenomenologically interesting region of magnetic fields for magnetars

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## References

- [1] B. W. Mintz, E. S. Fraga, G. Pagliara and J. Schaffner-Bielich, Phys. Rev. D **81**, 123012 (2010) [arXiv:0910.3927 [hep-ph]].
- [2] R. C. Duncan and C. Thompson, Astrophys. J. **392**, L9 (1992); C. Thompson and R. C. Duncan, *ibid.* **408**, 194 (1993).
- [3] O. Scavenius, A. Mocsy, I. N. Mishustin and D. H. Rischke, Phys. Rev. C **64**, 045202 (2001)
- [4] R. D. Pisarski and F. Wilczek, Phys. Rev. D **29**, 338 (1984).
- [5] O. Scavenius, A. Dumitru, E. S. Fraga, J. T. Lenaghan and A. D. Jackson, Phys. Rev. D **63**, 116003 (2001)
- [6] T. Fischer, S. C. Whitehouse, A. Mezzacappa, F. -K. Thielemann and M. Liebendorfer, Astronomy and Astrophysics **499**, 1 (2009)
- [7] I. Bombaci, D. Logoteta, P. K. Panda, C. Providencia and I. Vidana, Phys. Lett. B **680**, 448 (2009)
- [8] K. Iida and K. Sato, Prog. Theor. Phys. **98**, 277 (1997)
- [9] A. F. Garcia and M. B. Pinto, Phys. Rev. C **88**, no. 2, 025207 (2013)