

# Numerical holography and far-from-equilibrium dynamics

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# motivation

- Use gauge/gravity duality to study, far-from-equilibrium strongly interacting dynamics
- Go beyond near-equilibrium dynamics (linear response, probe approximation)
- Honestly solve dynamics of interesting initial states

# relativistic heavy ion collisions

Relevant dynamics:

Very early: partonic, perturbative (?)

Plasma phase: **strongly coupled**

Evidence: screening lengths, viscosity, ...

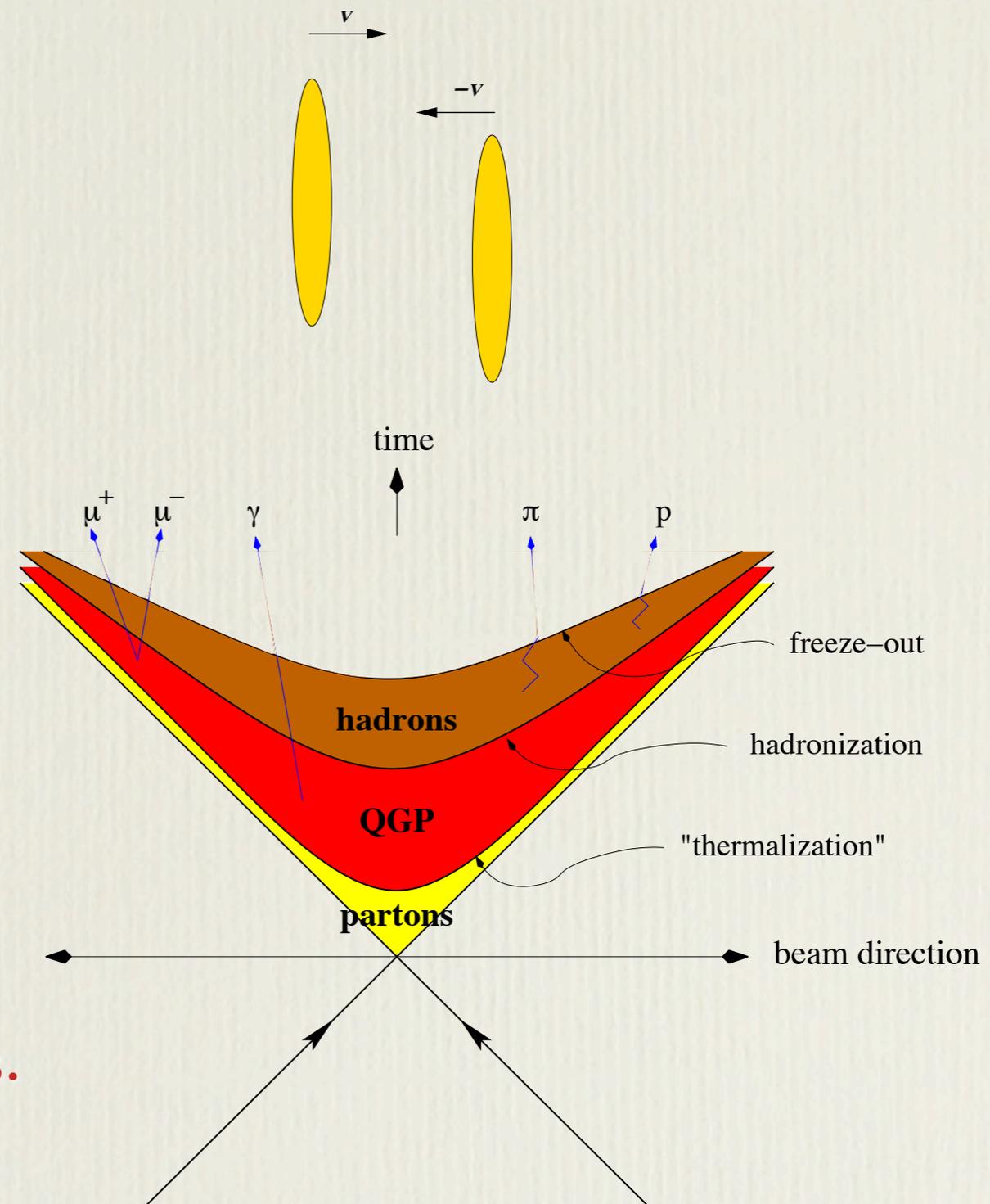
Many questions:

How fast do produced partons isotropize?

When/where is hydrodynamics valid?

Signatures of strongly coupled dynamics?

**No fully controlled theoretical methods.**



# relativistic heavy ion collisions

Idealize:

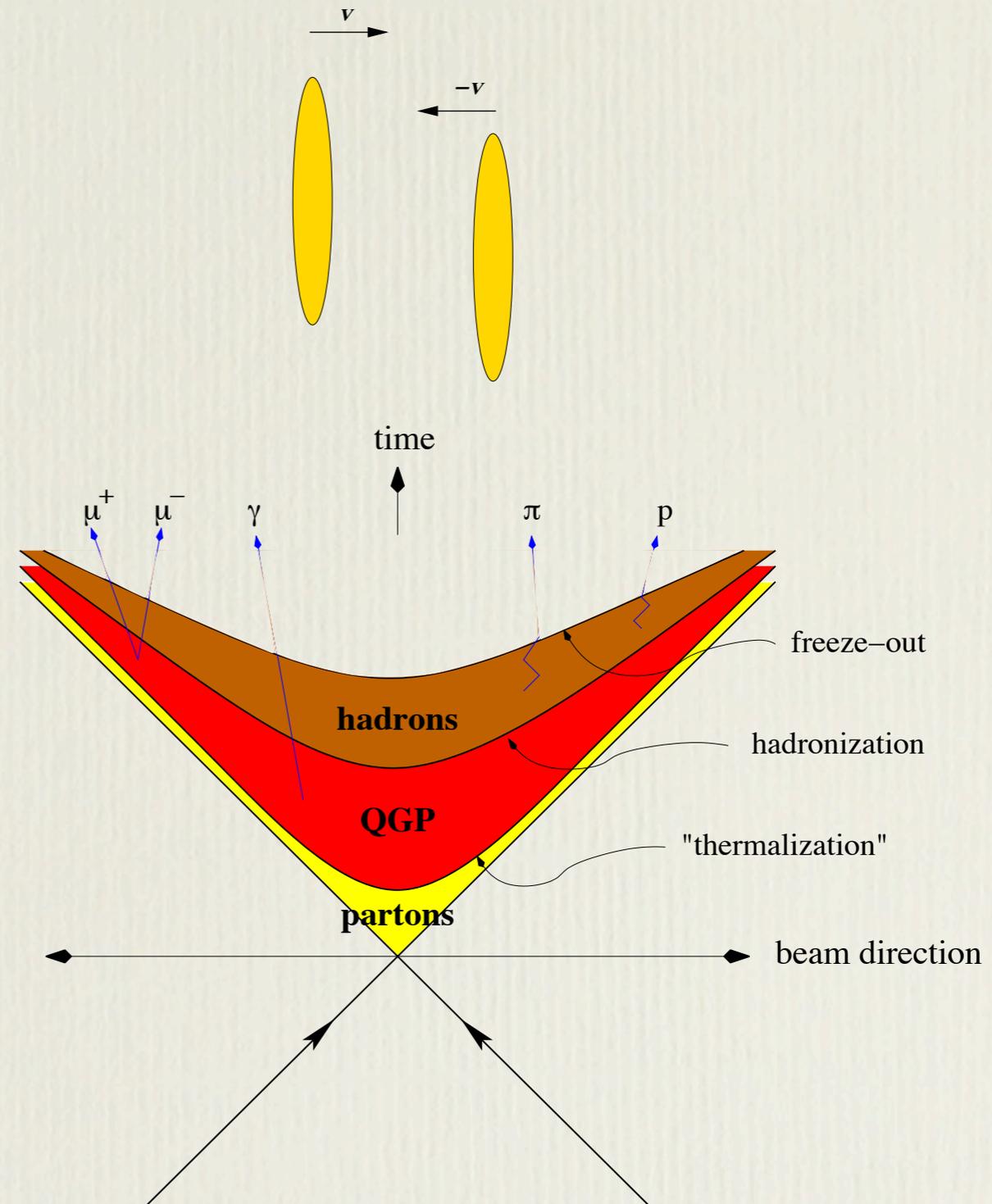
SU(3) gauge field + quarks  $\Rightarrow$   
SU( $N_c$ ) gauge field + adjoint matter

strongly coupled QCD  $\Rightarrow$   
strongly coupled  $\mathcal{N}=4$  SYM

# colors:  $N_c = 3 \Rightarrow N_c = \infty$

't Hooft coupling:  $\lambda \approx 1 \Rightarrow \lambda \gg 1$

Use holographic methods to study  
non-equilibrium, strongly coupled  
non-Abelian plasma dynamics

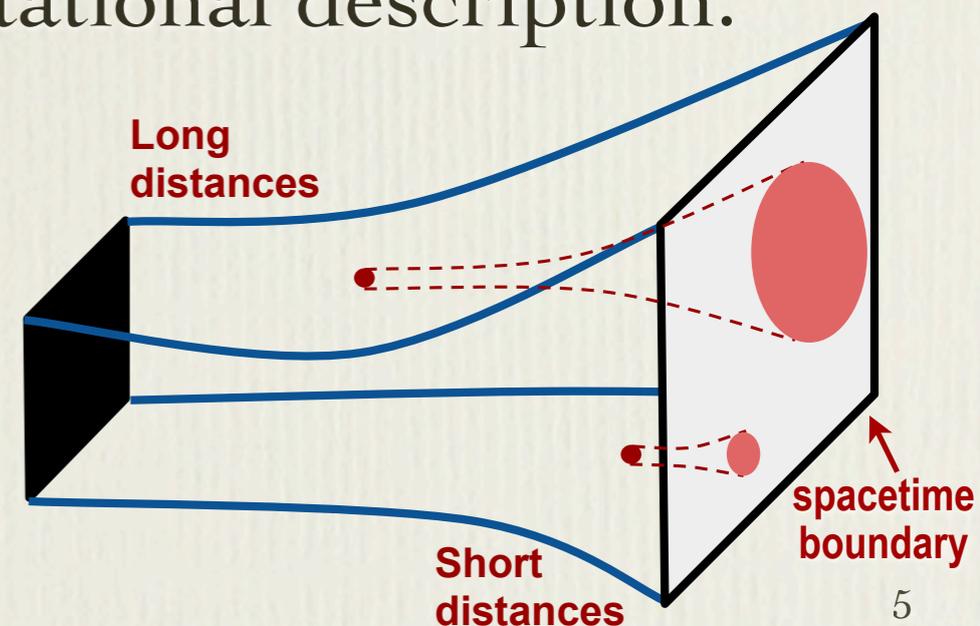


# gauge/gravity duality

- a.k.a. “AdS/CFT duality,” “gauge/string duality,” “holography”
- Some non-Abelian gauge theories have **exact** reformulation as higher dimensional gravitational (or string) theories.

Simplest case: maximally supersymmetric  $SU(N_c)$  Yang-Mills ( $\mathcal{N}=4$  SYM)  
= string theory on  $AdS_5 \times S^5$ . More complicated generalizations for less supersymmetric, non-conformal theories.

- Strong coupling (and large  $N_c$ ) limit of **quantum** field theory given by **classical** dynamics in dual gravitational description.
- Holographic description gives geometric representation of renormalization flow:



# holography: features

- strongly coupled, large  $N$  QFT = classical (super)gravity in higher dimension
  - valid description on all scales
  - gravitational fluctuations:  $1/N^2$  suppressed
  - QFT state  $\leftrightarrow$  asymptotically AdS geometry
  - $O(N^2)$  entropy  $\leftrightarrow$  gravitational (black brane) horizon
  - thermalization  $\leftrightarrow$  gravitational infall, horizon formation & equilibration

# applications of holography

- Equilibrium properties of strongly coupled theories:

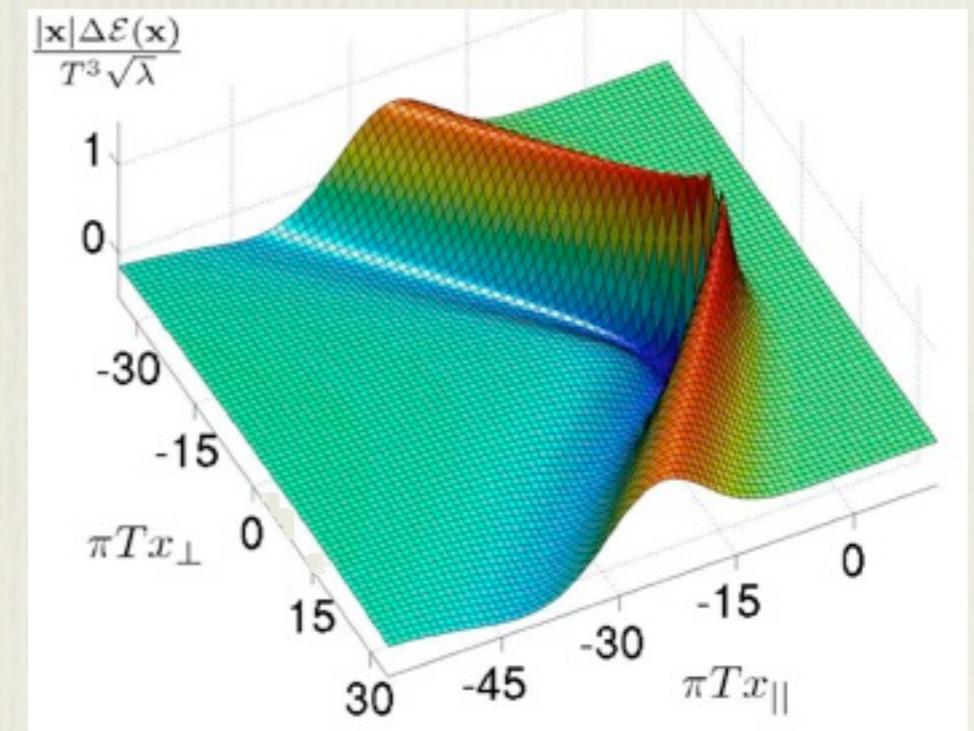
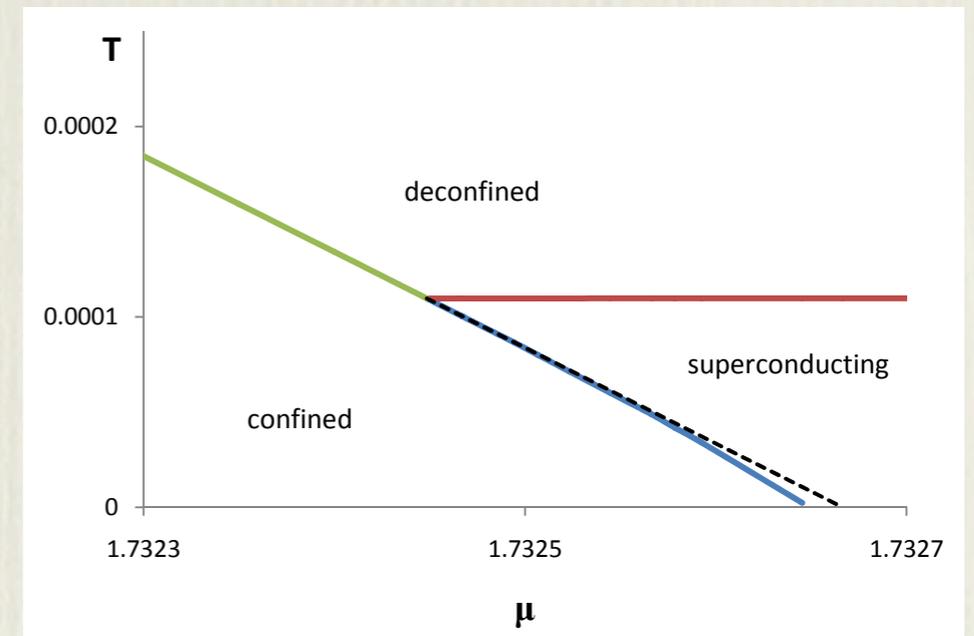
static geometries  $\Leftrightarrow$  equilibrium states of QFT  
AdS<sub>5</sub> black hole  $\Leftrightarrow$  non-Abelian plasma

- Near-equilibrium dynamics:

small fluctuations  $\Rightarrow$  linear response, spectral densities, transport coefficients, quasi-normal modes, photoemission, probe dynamics

- Far-from-equilibrium dynamics:

gravitational initial value problems



# far-from-equilibrium dynamics

- heavy-ion collisions:
  - homogeneous isotropization
  - boost invariant flow
  - colliding planar shocks
  - colliding “nuclei”
- turbulence:
  - 2D normal fluids
  - 2D superfluids
- other stuff:
  - dynamical quenches
  - black hole formation/ring-down

P. Chesler, L.Y., M. Heller, D. Mateos, W. van der Schee, ...

P. Chesler, L.Y., R. Janik, M. Heller, P. Witaszczyk, W. van der Schee

P. Chesler, L.Y., J. Casalderrey Solano, M. Heller, D. Mateos, W. van der Schee

W. van der Schee, P. Romatschke, S. Pratt, P. Chesler, L.Y.

A. Adams, P. Chesler, H. Liu

A. Adams, P. Chesler, H. Liu

M.J. Bhaseen, J. Gauntlett, J. Sonner, T. Wiseman, A. Garcia-Garcia; B. Craps, L. Lehner, K. Schalm, R. Myers

H. Bantilan, F. Pretorius, S. Gubser

# this talk

- Methods
- Colliding shocks
  - dependence on shock width?
  - surviving remnants of initial shocks?
  - approximate boost invariance?
  - finite size “nuclei”?
- Homogeneous equilibration
  - sensitivity to charge density or magnetic field?
  - degree of non-linearity?
- 2D turbulence
  - normal fluids
  - superfluids

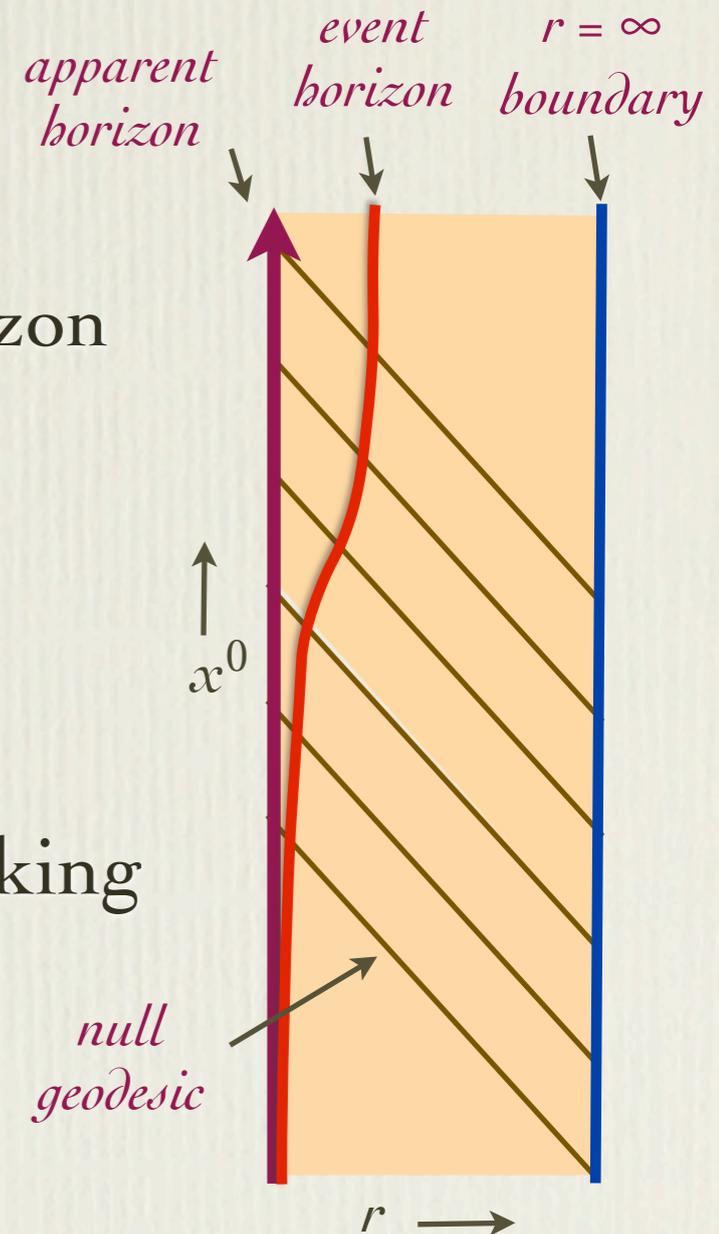
*work in progress  
w. Paul Chesler*

*work in progress  
w. John Fuini*

# methods

- Characteristic formulation

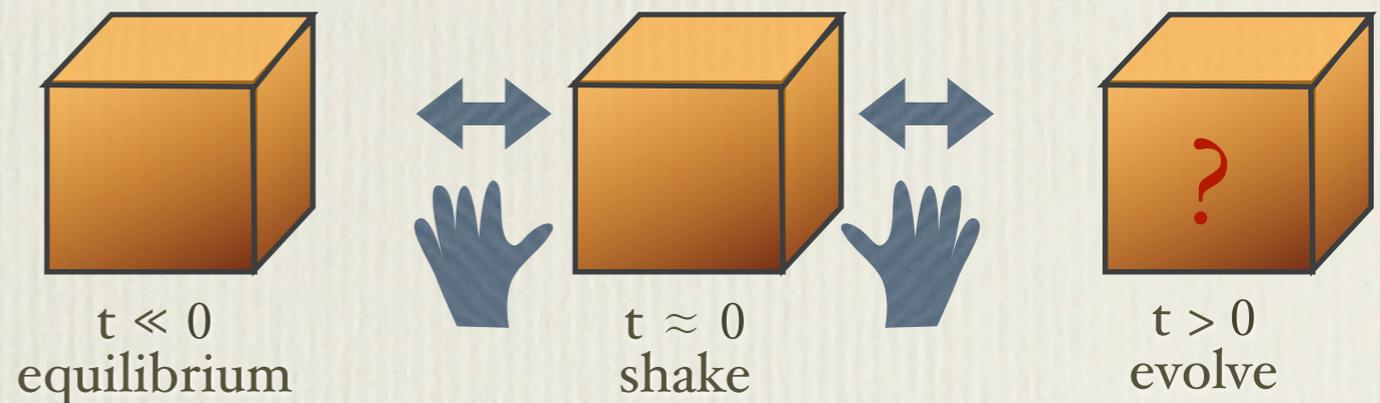
- coordinates adapted to infalling null geodesics
- fix residual diffeomorphisms: planar apparent horizon
- Einstein equations: nested linear radial ODEs
- discretize using pseudo-spectral derivatives
- low-pass filtering: alleviate aliasing, spectral blocking
- domain decomposition
- Fast, accurate, stable evolution achievable



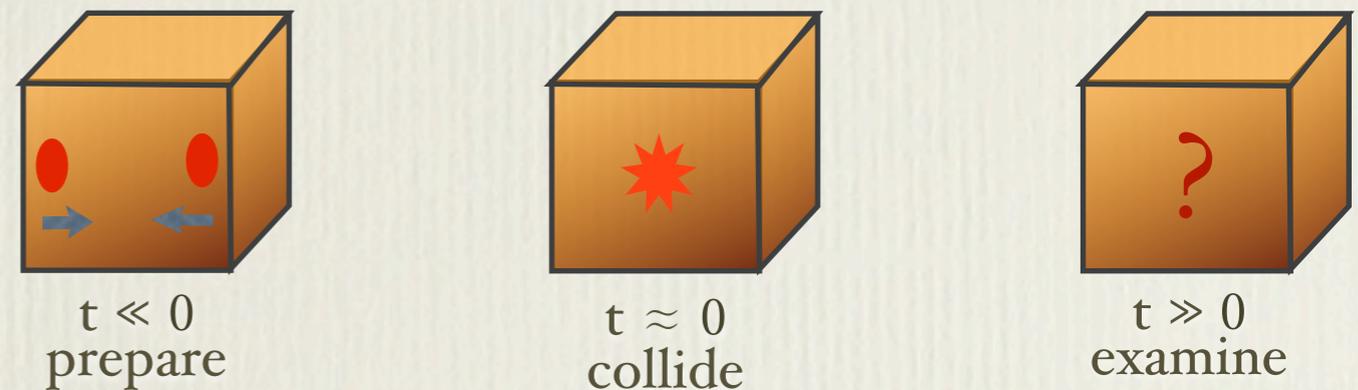
# initial data: choices

- Use time-dependent external fields:

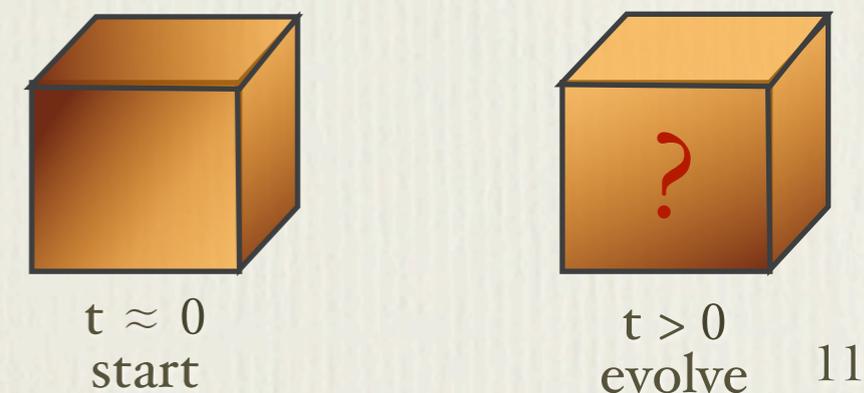
time-dependent dynamics  $\leftrightarrow$   
external work done on system



- Do scattering experiment:



- Choose geometry on initial (null) surface



# colliding planar shocks

- energy density localized on infinite planar sheets
- caricature of large, Lorentz-contracted nuclei
- questions:
  - domain of validity of hydrodynamic approximation?
  - dependence on longitudinal profile?
  - surviving remnants?
  - approximate boost invariance?



# colliding planar shocks

- 2D translation invariance  $\Rightarrow$  2+1D PDEs

$$ds^2 = \Sigma(X)^2 \hat{g}_{ij}(X) dx^i dx^j + 2dt [dr - A(X) dt - F_i(X) dx^i]$$

$$X = (t, z, r) \quad \|\hat{g}_{ij}\| = \text{diag}(e^B, e^B, e^{-2B})$$

- Initial conditions: superposition of counter-propagating planar shocks
  - Single shock, arbitrary longitudinal profile: known solution:

$$ds^2 = r^2[-dx_+ dx_- + dx_\perp^2] + \frac{1}{r^2} [dr^2 + h(x_\pm) dx_\pm^2]$$

Janik & Peschanski

- Choose Gaussian profile with width  $w$ , surface energy density  $\mu^3$ :

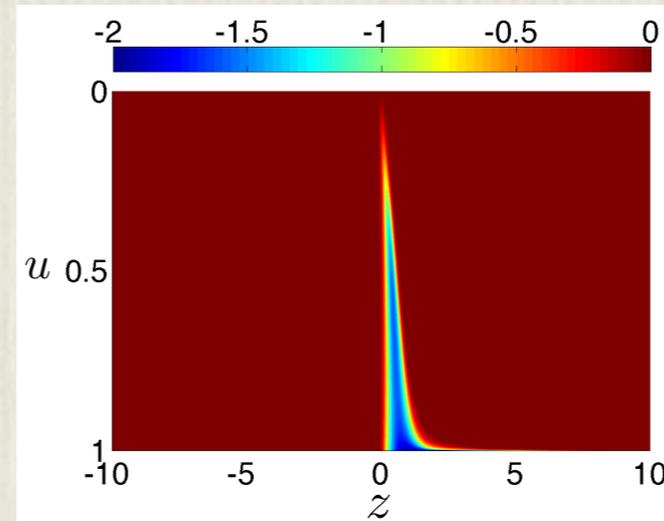
$$h(x_\pm) \equiv \mu^3 (2\pi w^2)^{-1/2} e^{-\frac{1}{2}x_\pm^2/w^2}$$

- Results depend on dimensionless width parameter  $w\mu$

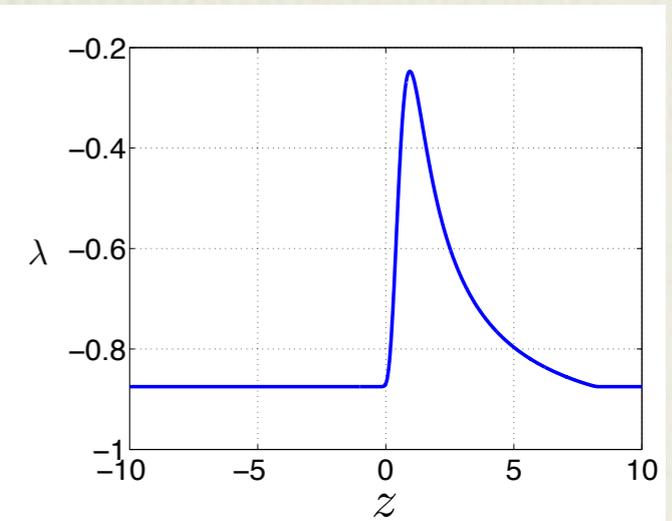
# initial data

- transformation to infalling coordinates:
  - must solve coupled 1+1D PDEs
  - shocks extend “forward” deep in bulk
  - apparent horizon exists regardless of separation

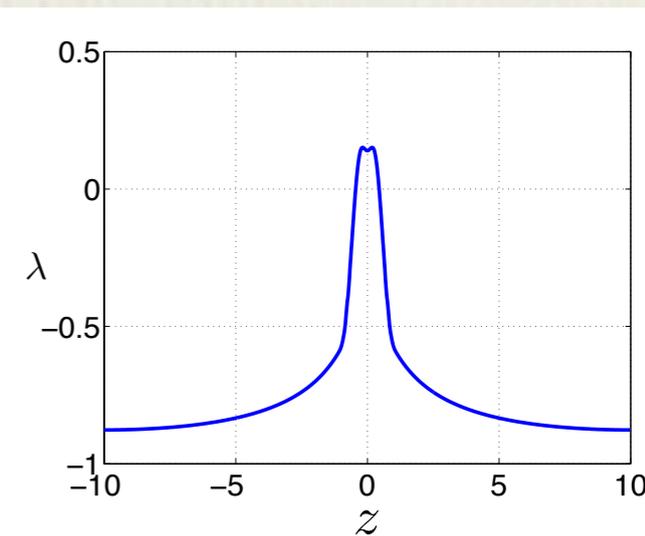
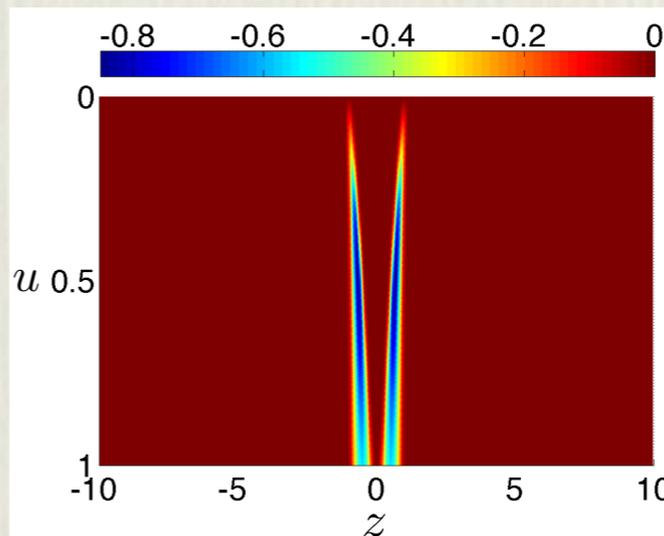
anisotropy  
function  $B(z,r)$



radial shift  
 $\lambda(z)$



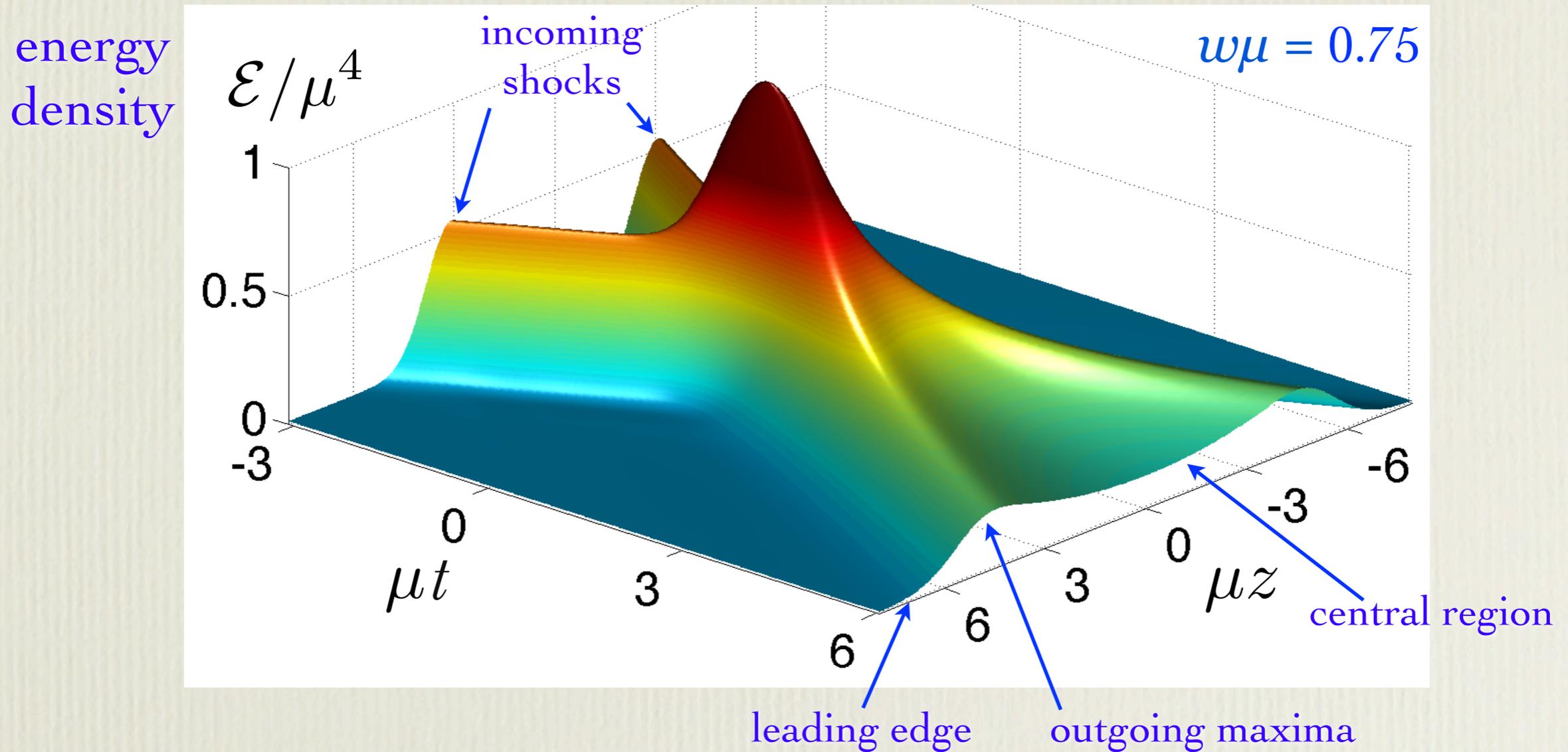
single right-moving shock



superposed shocks

# old results

C&Y: 1011.3562



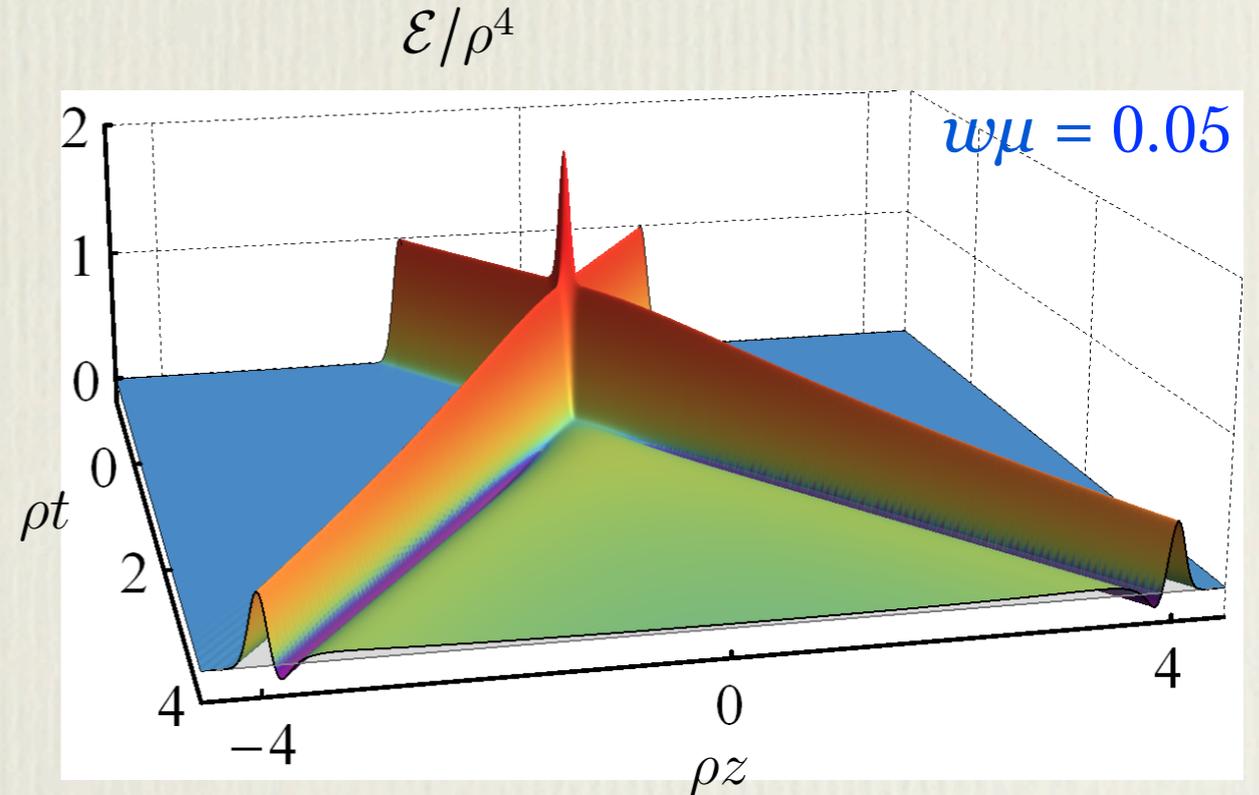
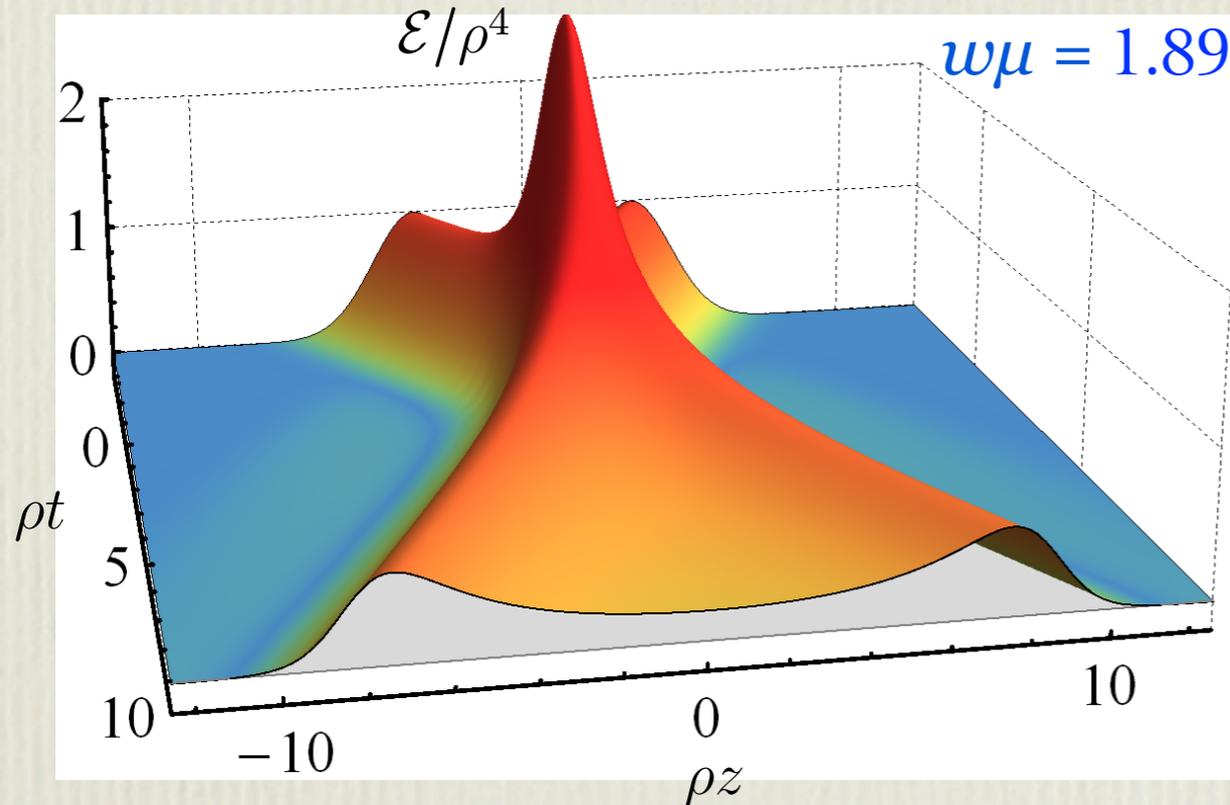
background energy density = 2% of single shock peak energy density

# more recent results

From full stopping to transparency in a holographic model of heavy ion collisions

Jorge Casalderrey-Solana,<sup>1</sup> Michal P. Heller,<sup>2,\*</sup> David Mateos,<sup>3,4</sup> and Wilke van der Schee<sup>5</sup>

1305.4919



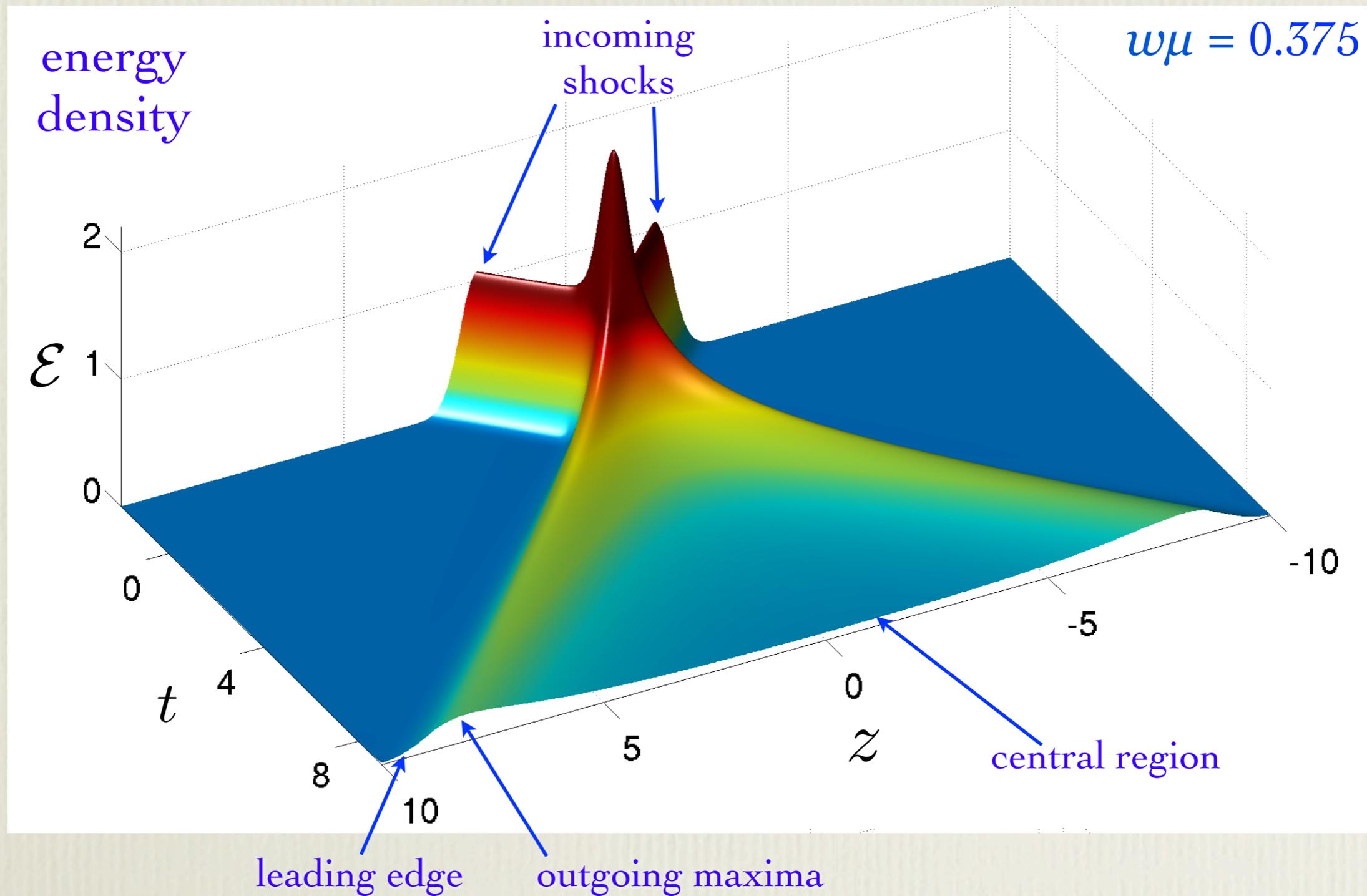
background energy density = 1.5 - 7.5% of single shock peak energy density

“We uncover a cross-over between two different dynamical regimes...  
At high energies, receding fragments move outward at the speed of light.”

# newer results (I)

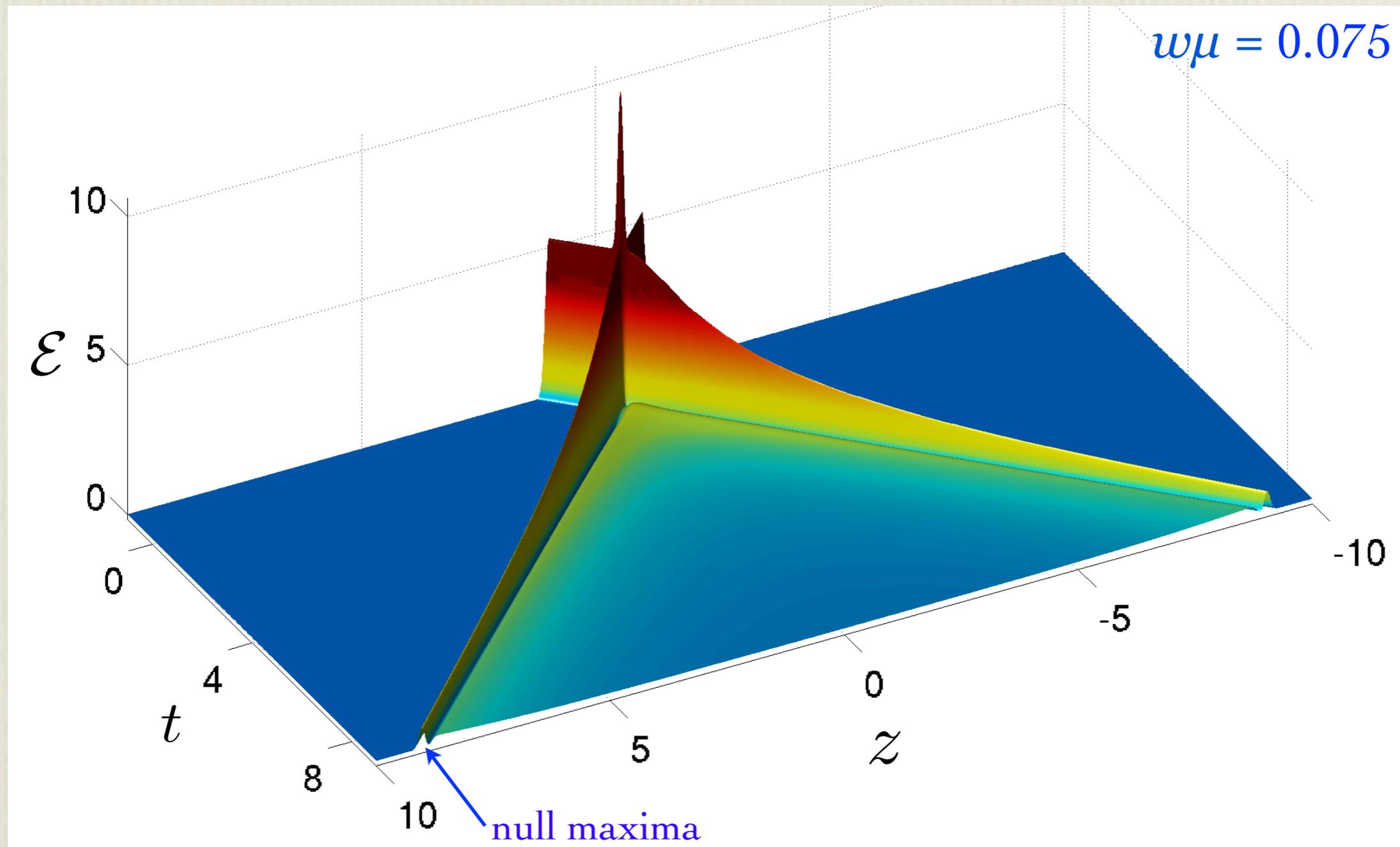
C&Y: 1309.1439

New: no background energy density

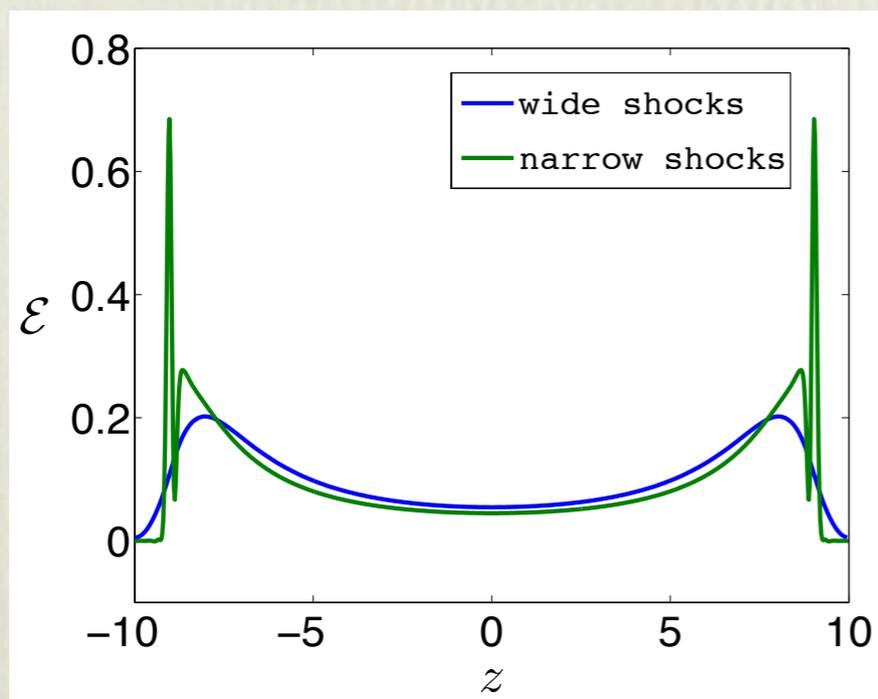
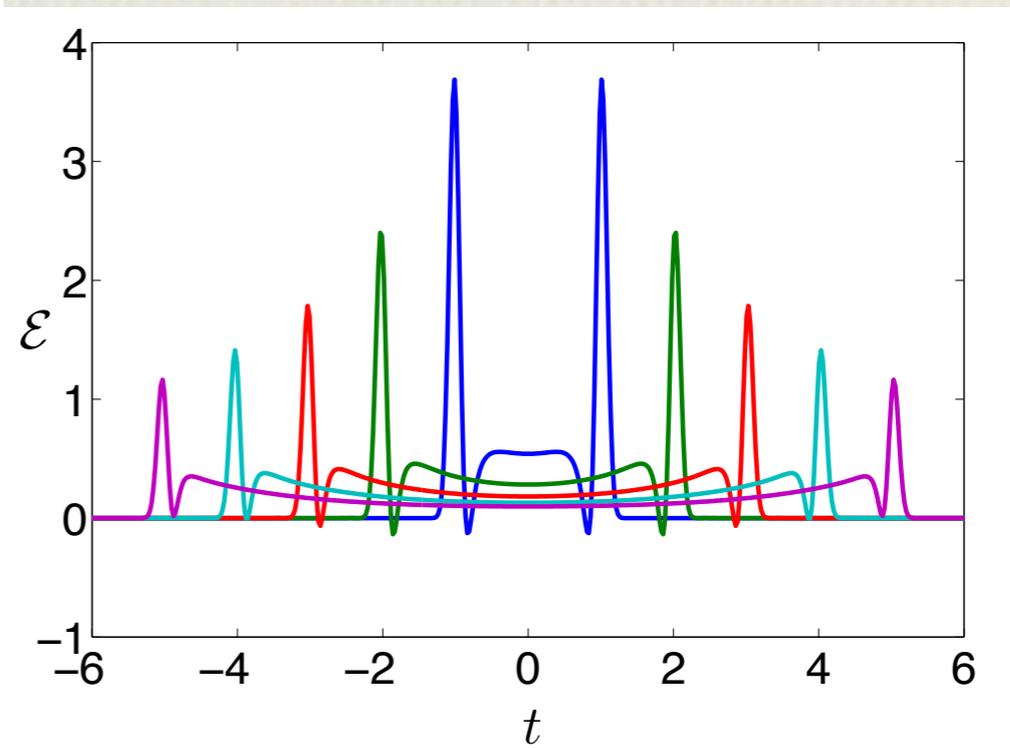
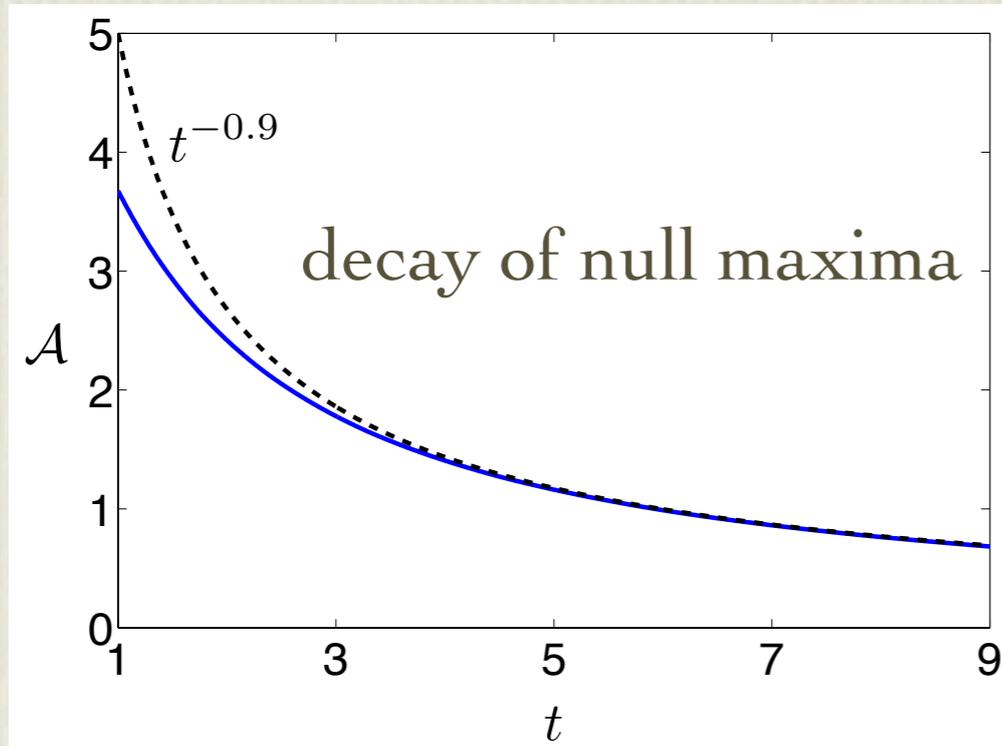


# newer results (II)

no background energy density,  
longer time evolution



# qualitative features



wide vs. narrow shocks,  
 $t = 9$

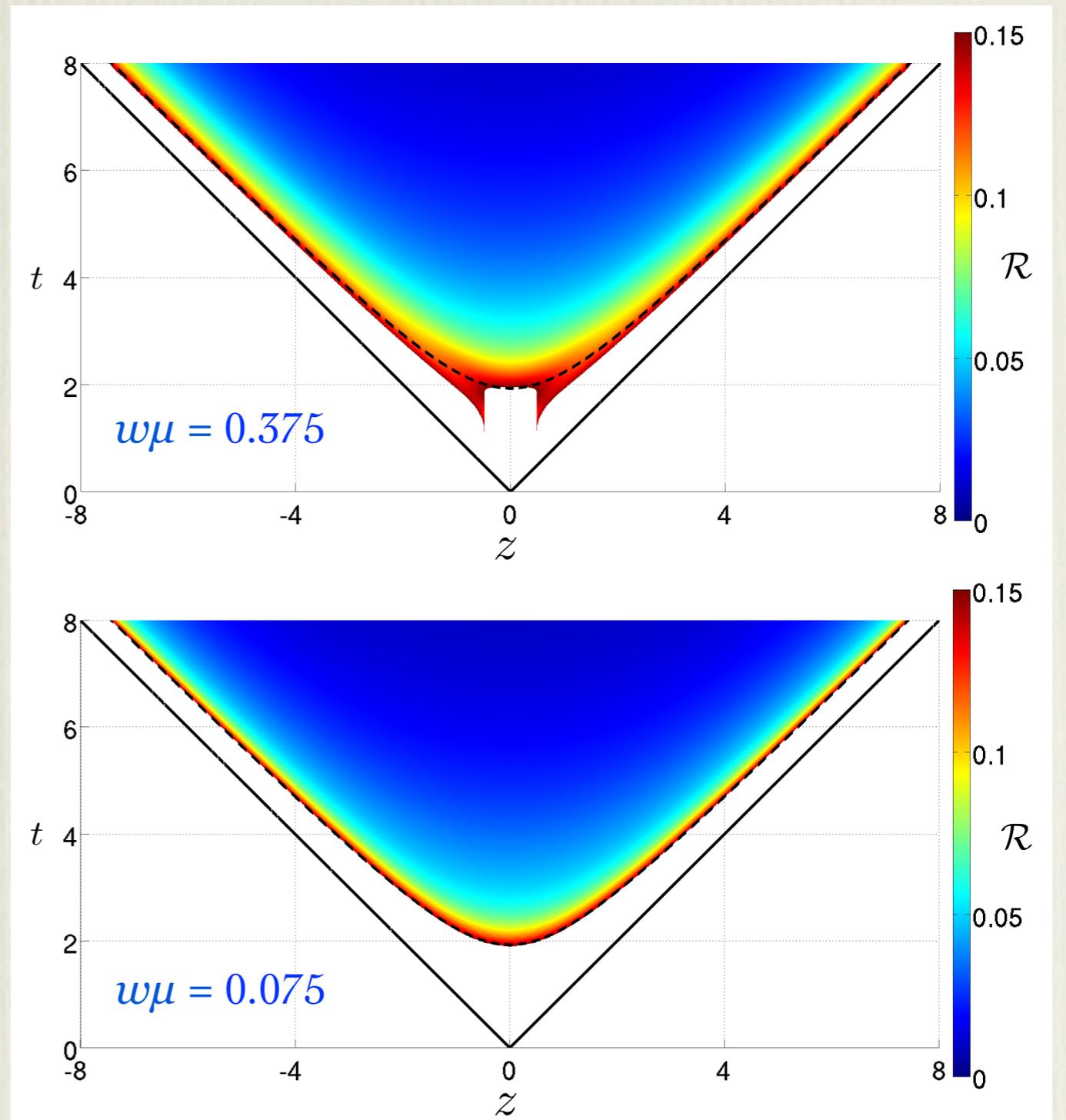
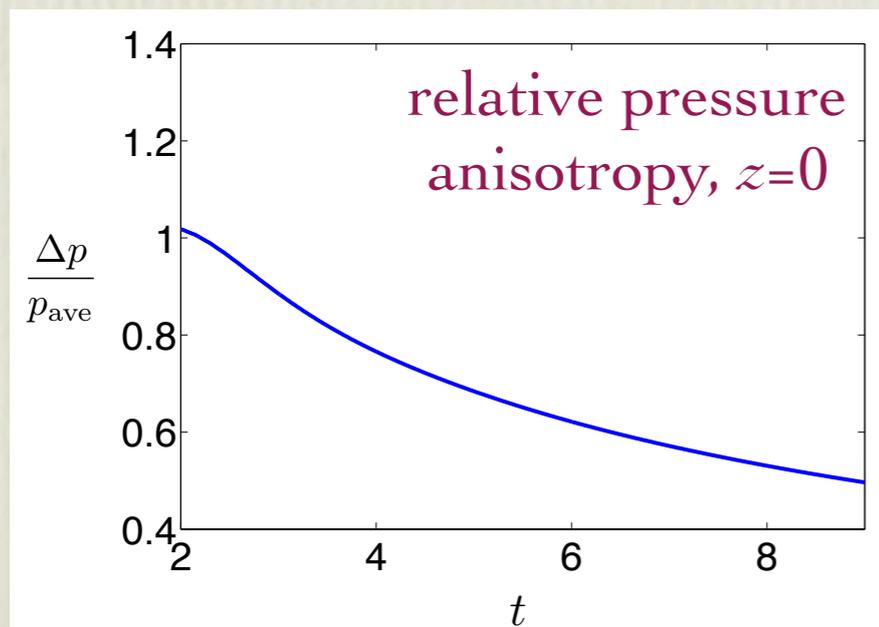
# validity of hydrodynamics

figure of merit:

$$\mathcal{R} \equiv \frac{1}{\bar{p}} \left[ (\langle T^{xx} \rangle - T_{\text{hydro}}^{xx})^2 + (\langle T^{zz} \rangle - T_{\text{hydro}}^{zz})^2 \right]^{1/2} \leq 15\%$$

↑  
first order viscous

hydro works even when  
viscous effects are  $O(1)$ :



# local boost invariance

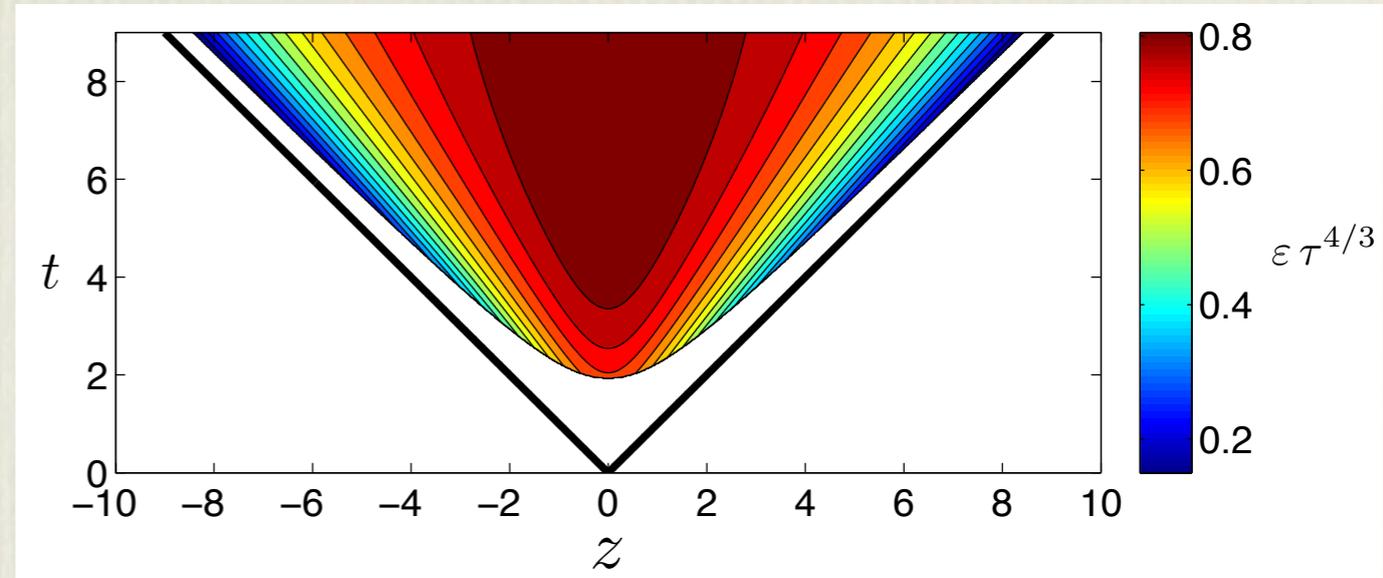
boost invariant flow:

$$u_\mu dx^\mu = d\tau \equiv \cosh y dt + \sinh y dz,$$

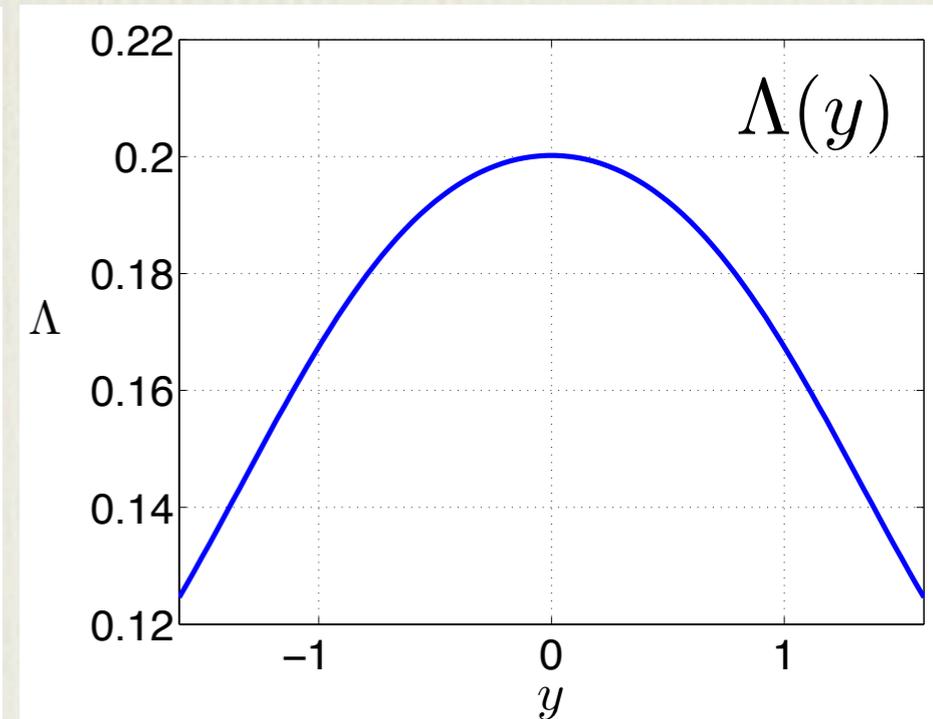
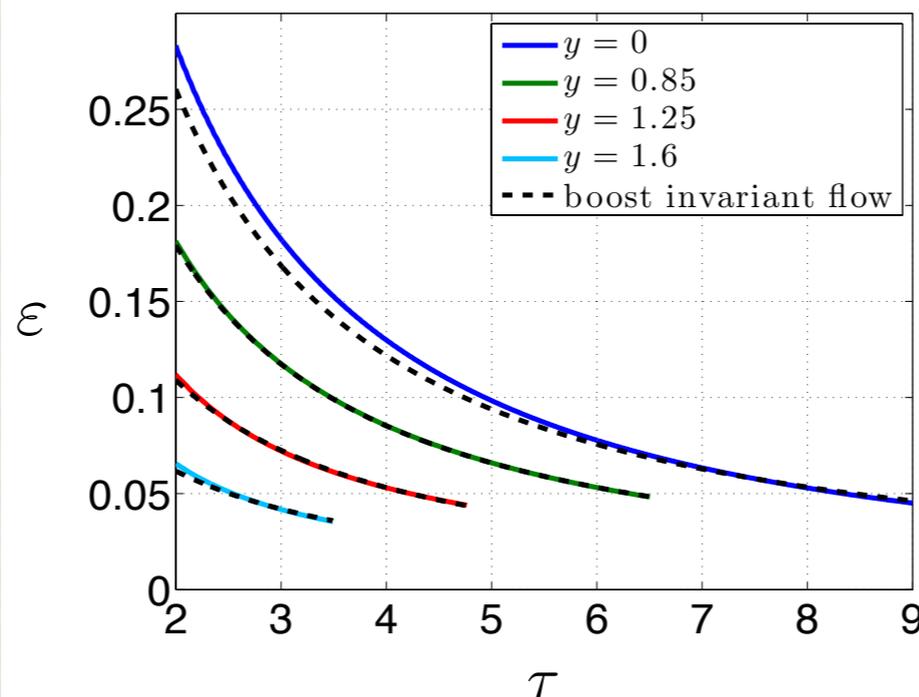
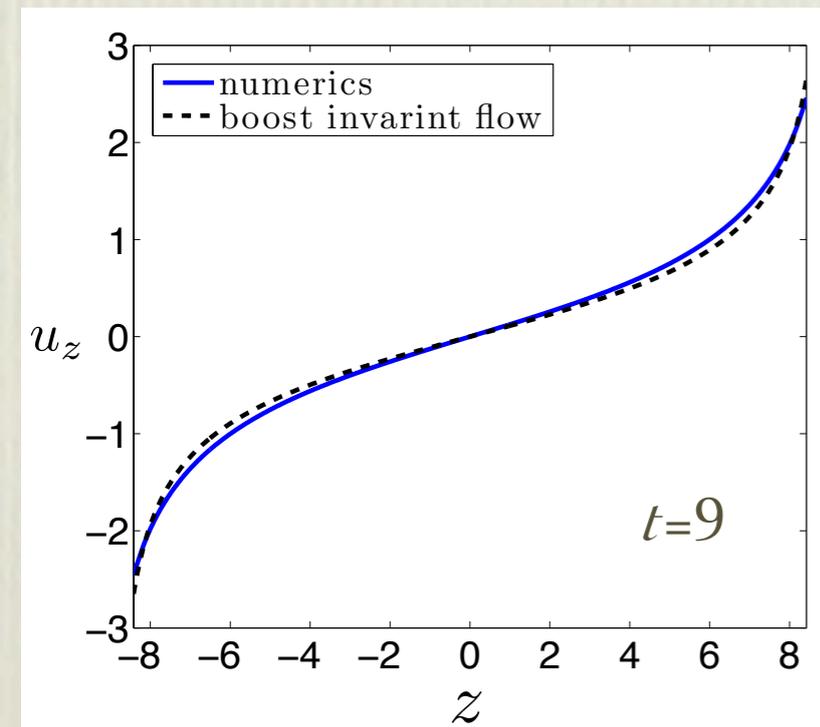
$$\varepsilon = \frac{3}{4} \frac{(\pi\Lambda)^4}{(\Lambda\tau)^{4/3}} \left[ 1 - \frac{C_1}{(\Lambda\tau)^{2/3}} + \frac{C_2}{(\Lambda\tau)^{4/3}} + O\left(\frac{1}{(\Lambda\tau)^2}\right) \right]$$

proper time  $\tau \equiv \sqrt{t^2 - z^2}$

rapidity  $y \equiv \tanh^{-1}(z/t)$

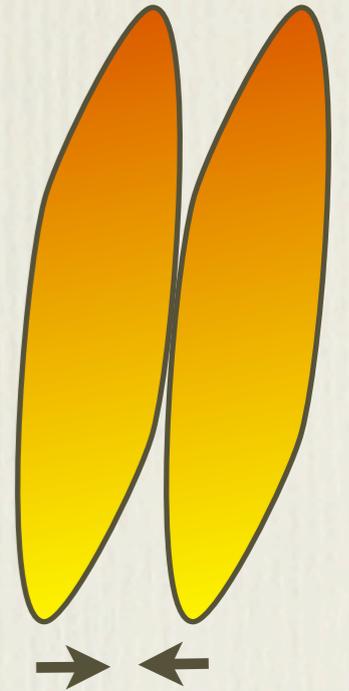


local boost invariance:  $\Lambda \rightarrow \Lambda(y)$



# colliding “nuclei”

- finite transverse extent, cylindrically symmetric
- single “nucleus”: smooth, localized null “shock”
  - ✓ exact solution = linear superposition of infinitely boosted point sources
  - ✓ transformation to null infalling coordinates
- implementation for general 4+1D case: in progress



Gubser, Pufu, Yarom

# colliding shocks: lessons

- Early times: large anisotropy, far from local equilibrium
- Rapidly attenuating outgoing maxima, no surviving remnants
- Mid-rapidity: hydrodynamics quickly becomes valid, despite large viscous effects

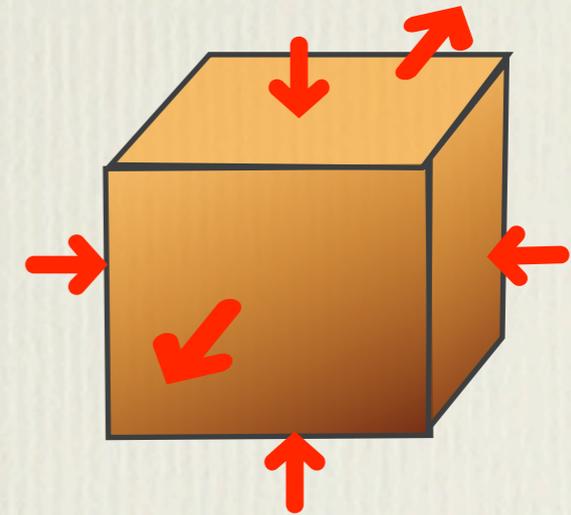
onset of hydro validity  $\approx 4/\mu$  after initial interaction

$\mu \approx 2.3$  GeV for modeling RHIC  $\Rightarrow \tau_{\text{hydro}} \approx 0.35$  fm/c

- Near outgoing lightcone: hydrodynamics not reliable
- Central region: “local” but not global boost-invariance

# homogeneous isotropization

- caricature of early moments of QGP
- no spatial gradients
  - ➔ no hydrodynamic response
  - ➔ exponential relaxation
- questions:
  - relaxation time scale?
  - onset of linearized regime?
  - sensitivity to (baryon) charge density?
  - sensitivity to magnetic field?



# homogeneous isotropization

arXiv:0812.2053

spatial homogeneity

$$\Rightarrow g_{\mu\nu} = g_{\mu\nu}(t, r)$$

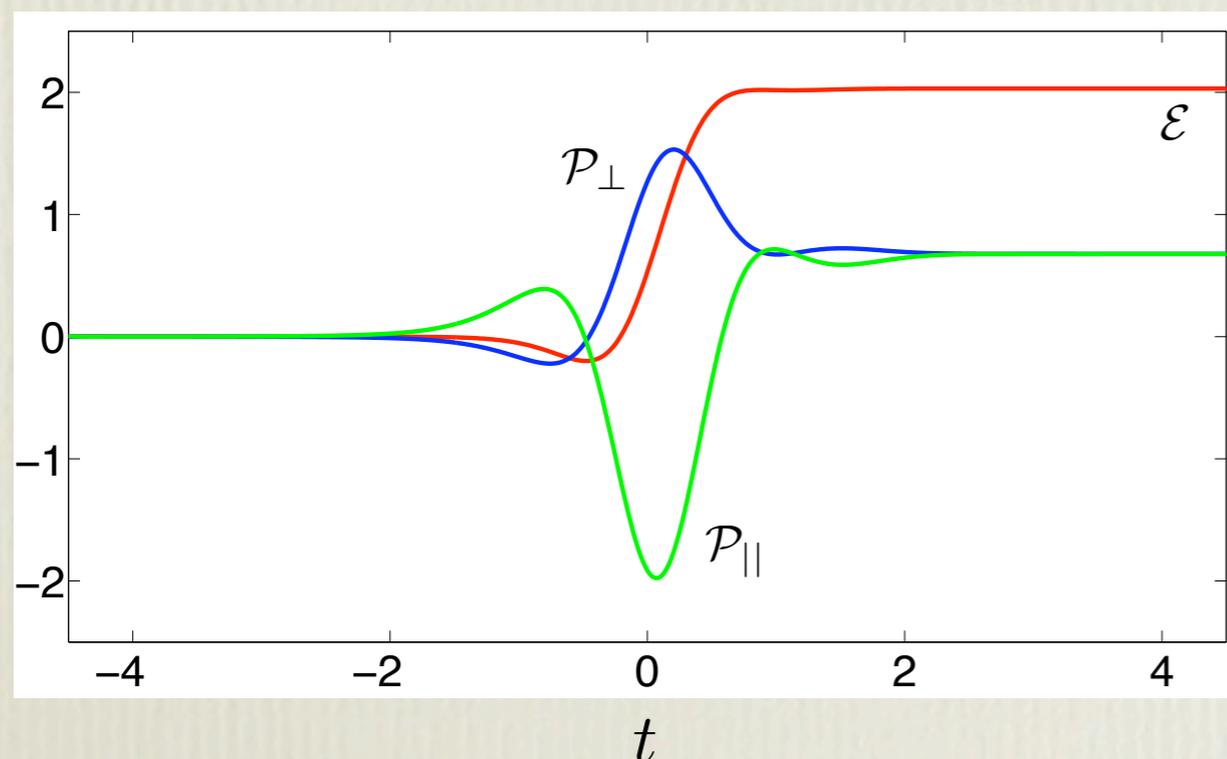
$\Rightarrow$  1+1D PDEs

$$\|\hat{g}_{ij}\| = \text{diag}(e^B, e^B, e^{-2B})$$

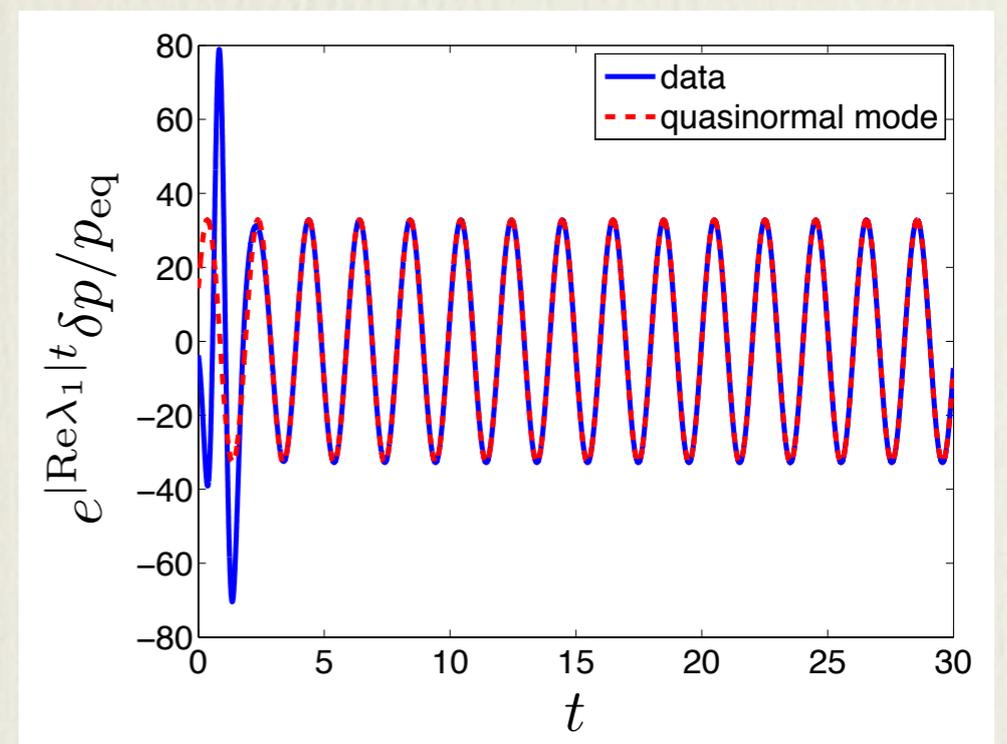
$$\begin{aligned} 0 &= \Sigma'' + \frac{1}{2}B'^2 \Sigma \\ 0 &= \Sigma (\dot{\Sigma})' + 2\Sigma' \dot{\Sigma} - 2\Sigma^2, \\ 0 &= \Sigma (\dot{B})' + \frac{3}{2}(\Sigma' \dot{B} + B' \dot{\Sigma}), \\ 0 &= A'' + 3B' \dot{B} - 12\Sigma' \dot{\Sigma}/\Sigma^2 + 4. \end{aligned}$$

$$\dot{h} \equiv d_+ h, \quad h' \equiv \partial_r h$$

transverse & longitudinal pressure



(rescaled) pressure anisotropy vs. lowest quasinormal mode



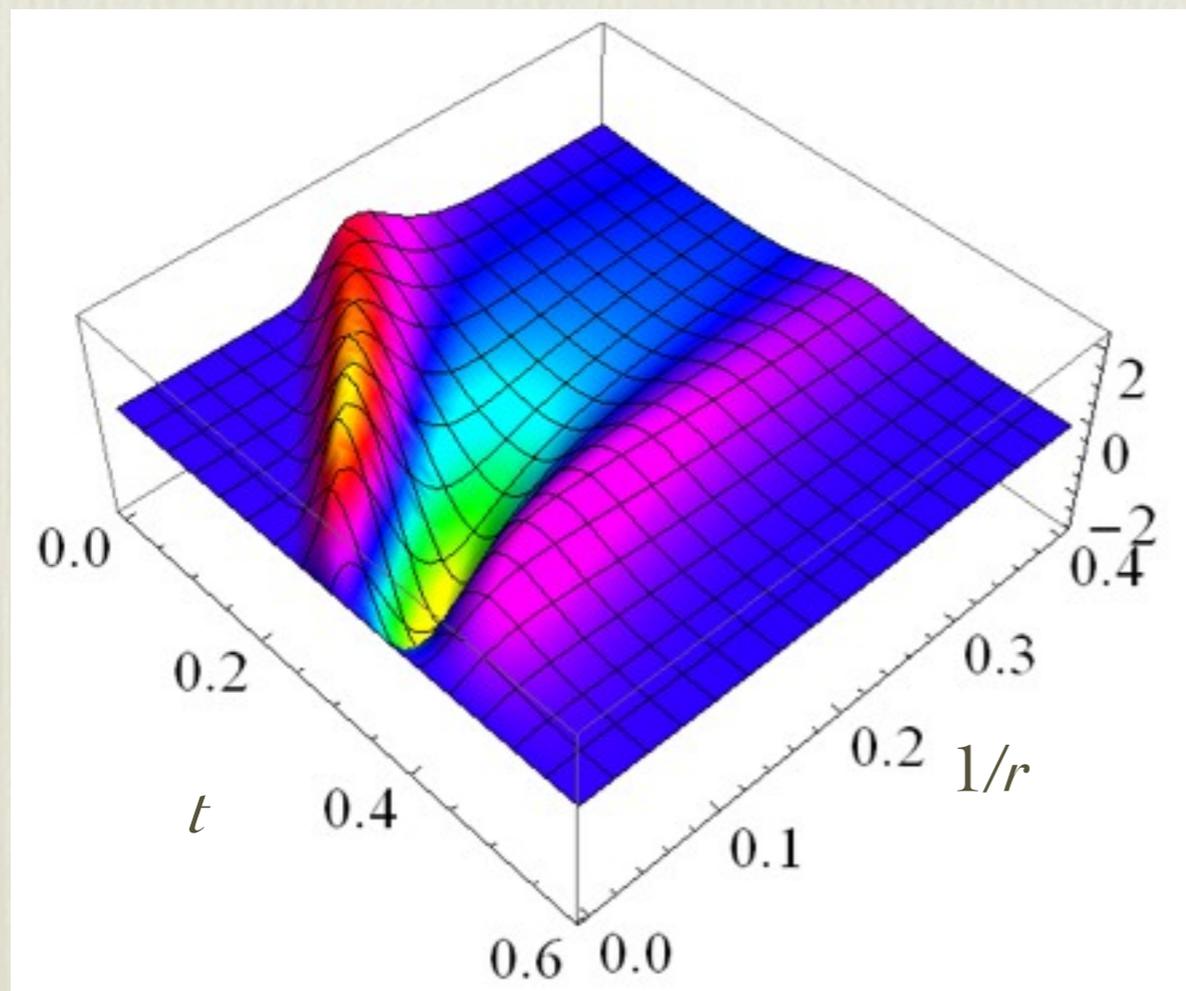
# homogeneous isotropization: recent work

- sensitivity to initial data choice
- sensitivity to non-zero charge density
- sensitivity to non-zero magnetic field
- deviation from linearized dynamics

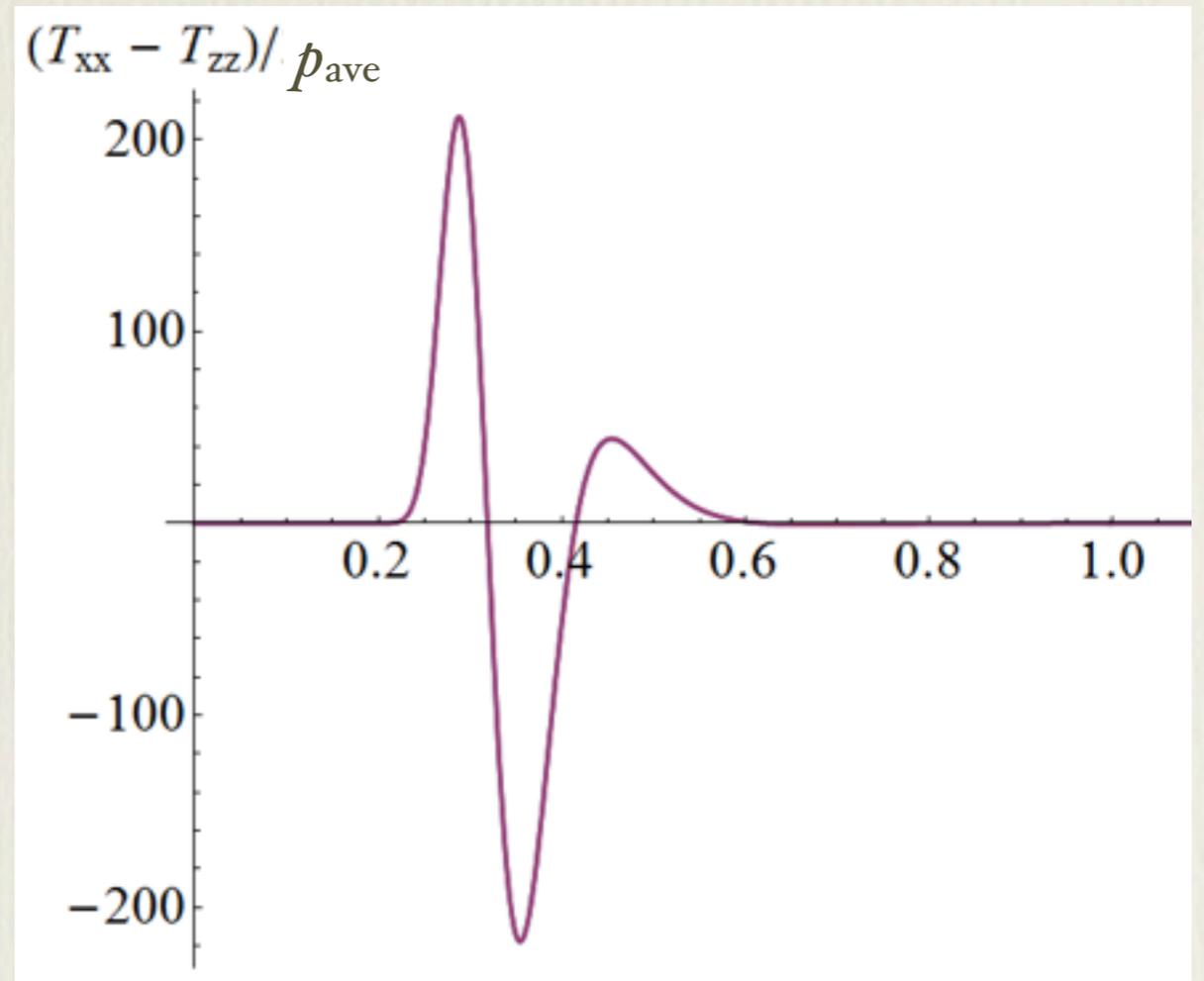
# homogeneous isotropization

initial data: Gaussian anisotropy function  $B(r)$

anisotropy function  $B(r)$



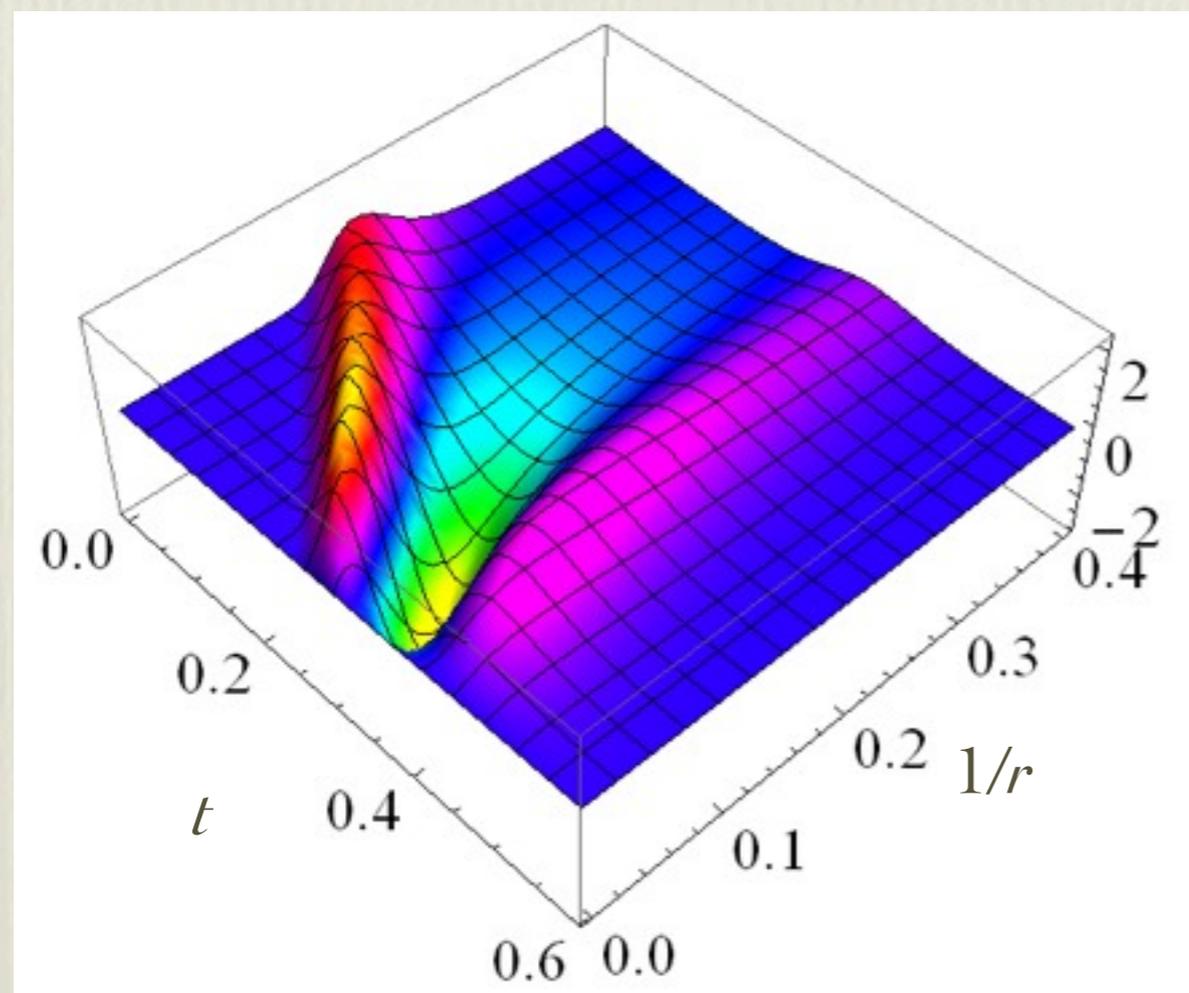
transverse & longitudinal pressure



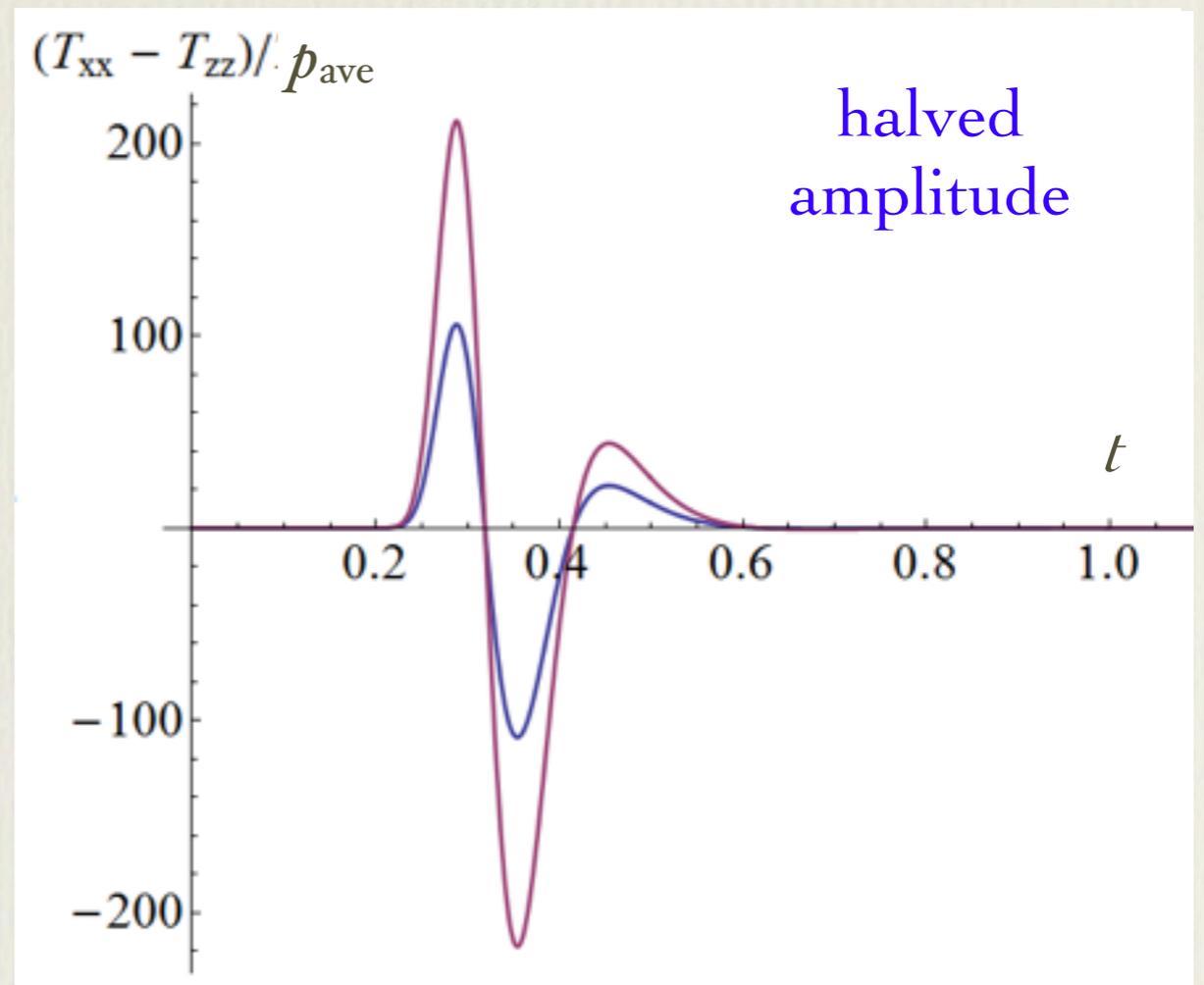
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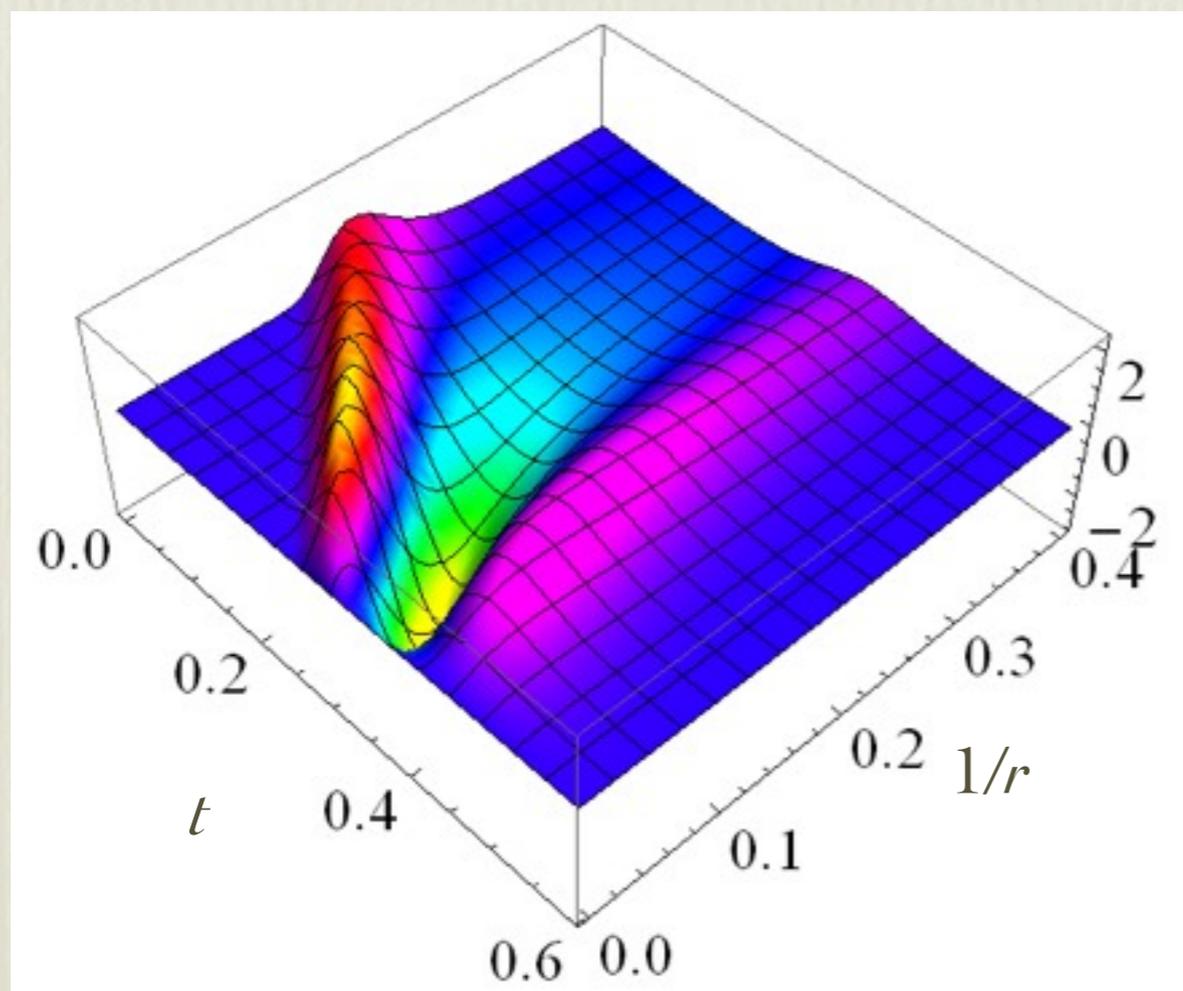
transverse & longitudinal pressure



# homogeneous isotropization

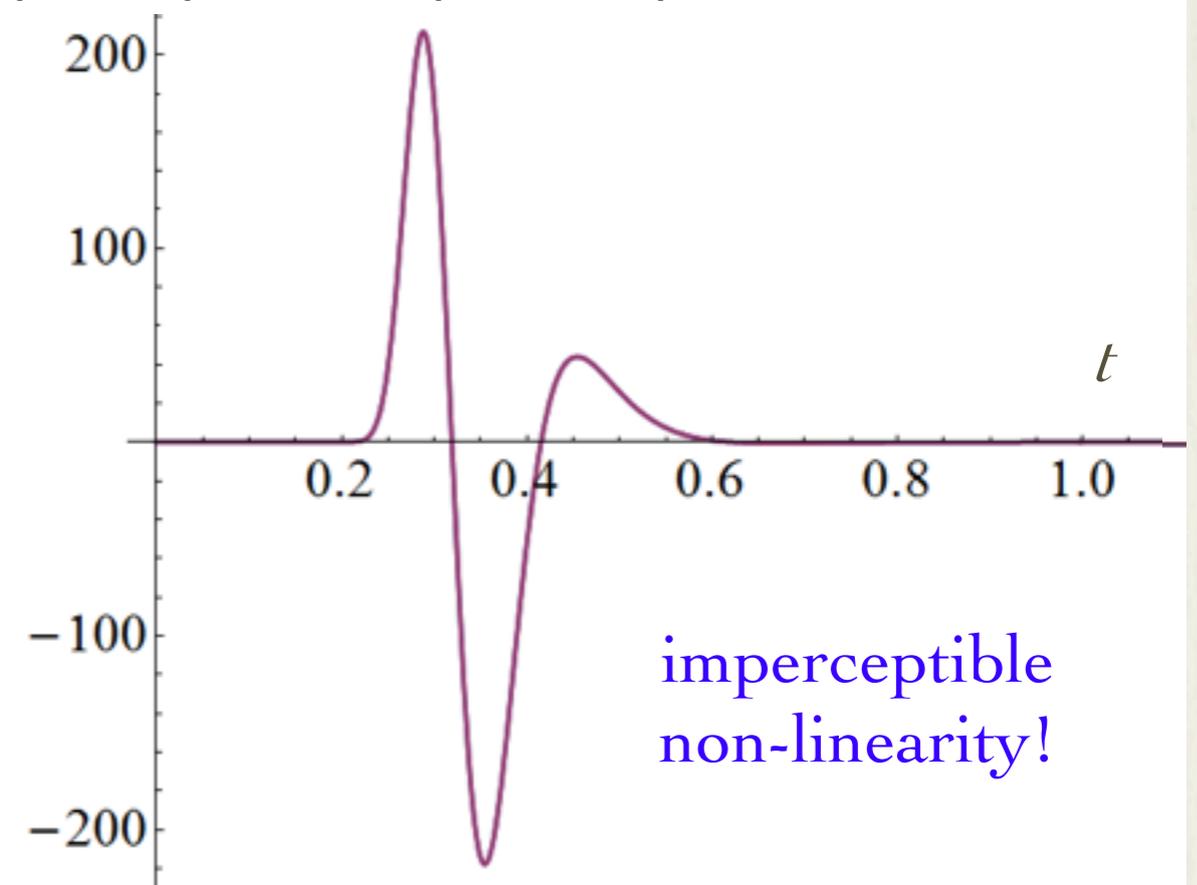
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transverse & longitudinal pressure

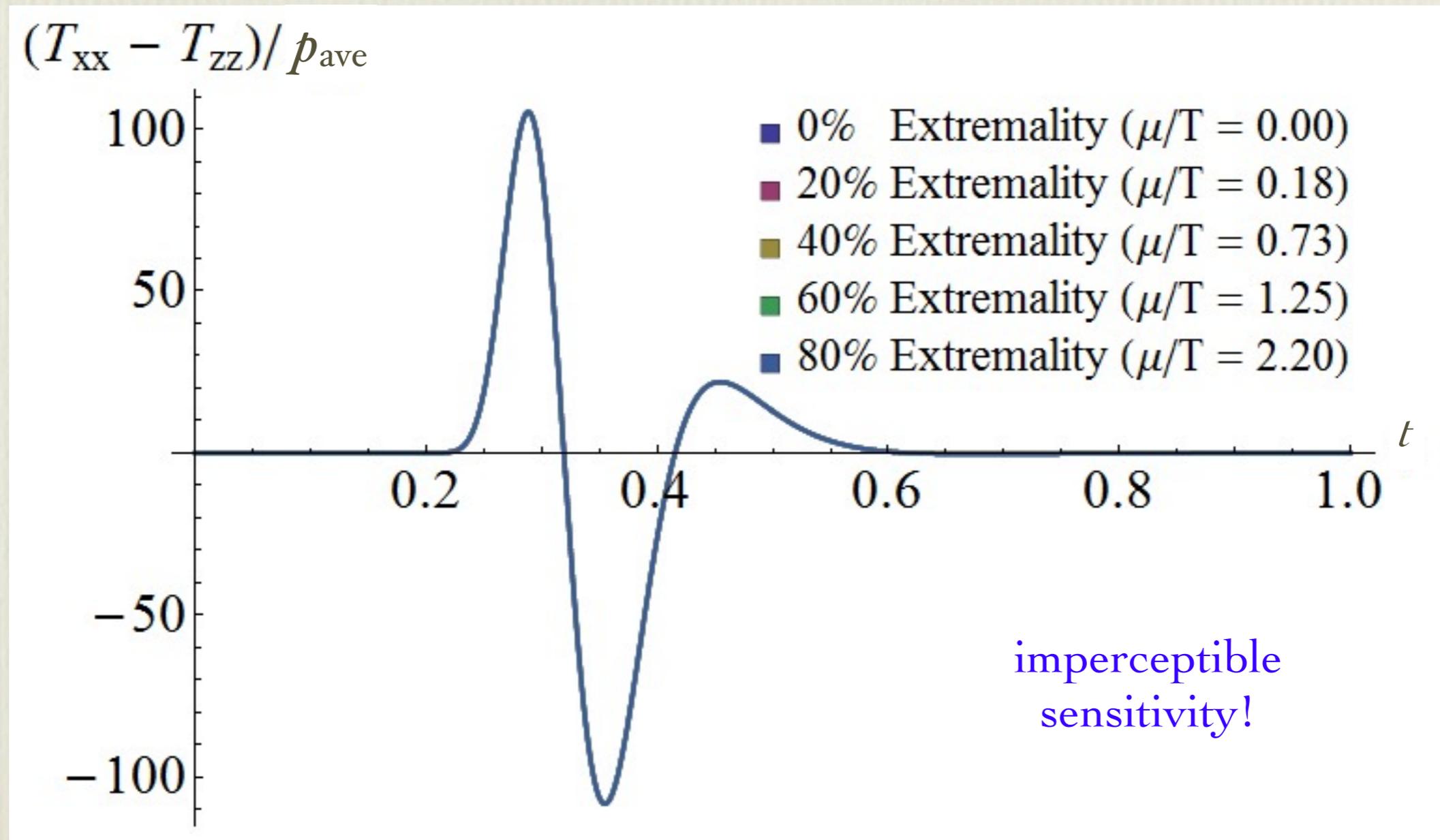
$\delta p(B_0)/\rho_{\text{ave}}$  &  $2\delta p(B_0/2)/\rho_{\text{ave}}$



# non-zero charge density

- heavy ion collisions:  $\mu/T = O(0.2)$
- sensitivity to plasma constituents?
- equilibrium state  $\leftrightarrow$  Reissner-Nordstrom black brane
  - maximal charge density  $\leftrightarrow$  extremal brane
  - $O(10\%)$  change in quasi-normal mode frequencies

# non-zero charge density



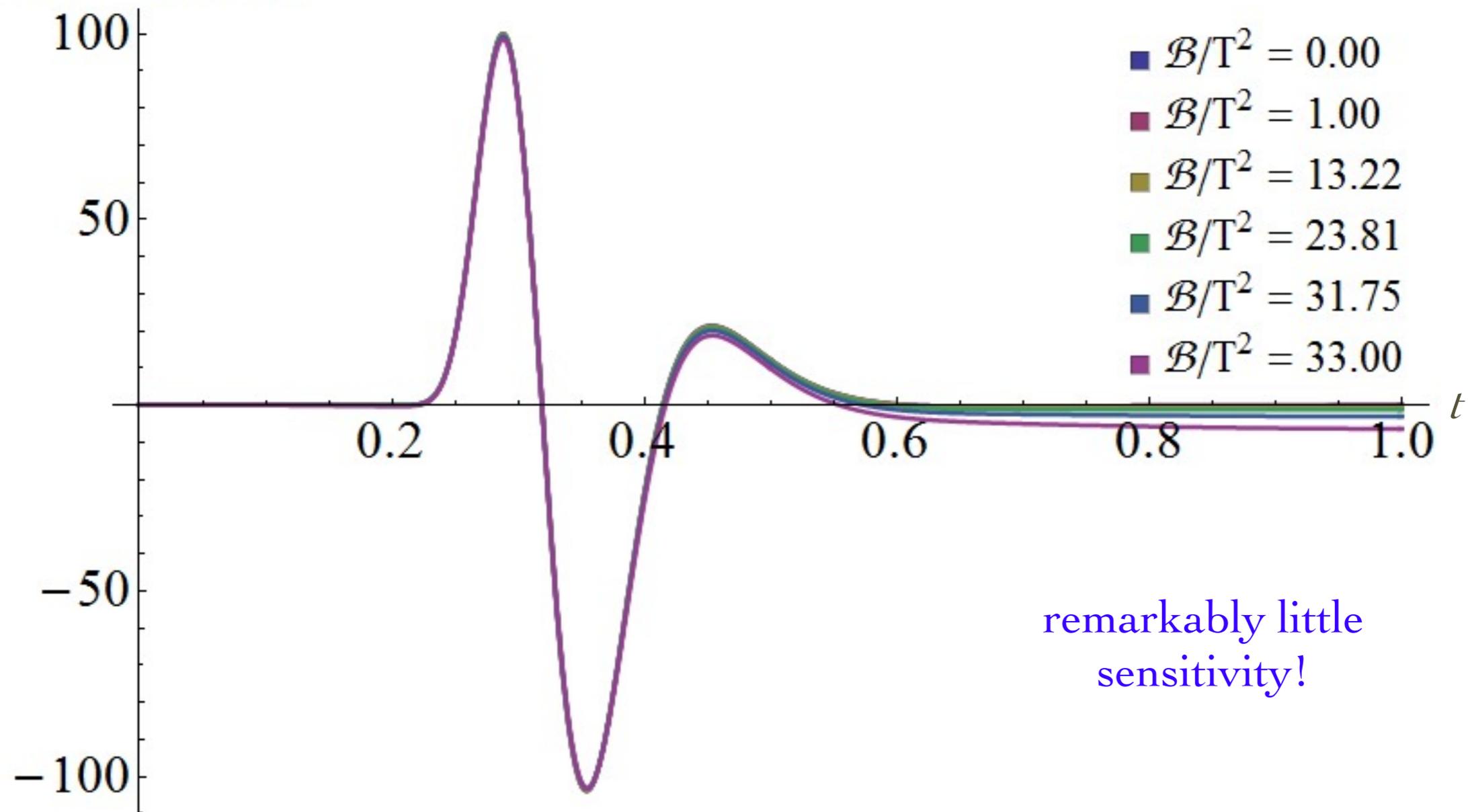
# non-zero magnetic field

- heavy ion collisions: much concern about E&M effects
- sensitivity to magnetic field?
  - externally imposed anisotropy
  - trace anomaly  $\Rightarrow$  broken scale invariance
  - $\langle T_{\mu\nu} \rangle = \kappa \left[ \tilde{g}_{\mu\nu}^{(4)} - \eta_{\mu\nu} \text{tr}(\tilde{g}^{(4)}) + c_0 \tilde{h}_{\mu\nu}^{(4)} \right]$   $\|\tilde{h}_{\alpha\beta}^{(4)}\| = \frac{1}{2} \mathcal{B}^2 \text{diag}(+1, +1, +1, -1)$

 classical E&M  
stress-energy

# non-zero magnetic field

$(T_{xx} - T_{zz})/T_{00}$  (static contribution omitted)



# homogeneous isotropization: lessons

- relaxation time scale = gravitational infall time
- remarkably little sensitivity to plasma constituents
- remarkably little sensitivity to added magnetic field
- dynamics, as probed by boundary observables, is nearly linear even very, very far from equilibrium!

# remarks (I)

- using gauge/gravity duality to study strongly coupled far-from-equilibrium dynamics **works** for interesting variety of problems
  - characteristic formulation, adapted to gravitational infall ➡ remarkably simple equations allowing efficient integration
  - can achieve stable evolution
  - desktop resources suffice for 1+1D, 2+1D, and even 3+1D problems
- no need to be professional numerical relativist!

# remarks (II)

- work to date has only scratched the surface; many interesting generalizations await:
  - collisions:
    - asymmetric planar shocks
    - “nuclei” with finite transverse extent
  - turbulence in three spatial dimensions:
    - normal fluids
    - superfluids
  - dynamics in non-conformal theories with more complicated dual gravitational descriptions